

Photographing Black Holes with the Event Horizon Telescope

- Past, Present and Future -



Kazu Akiyama

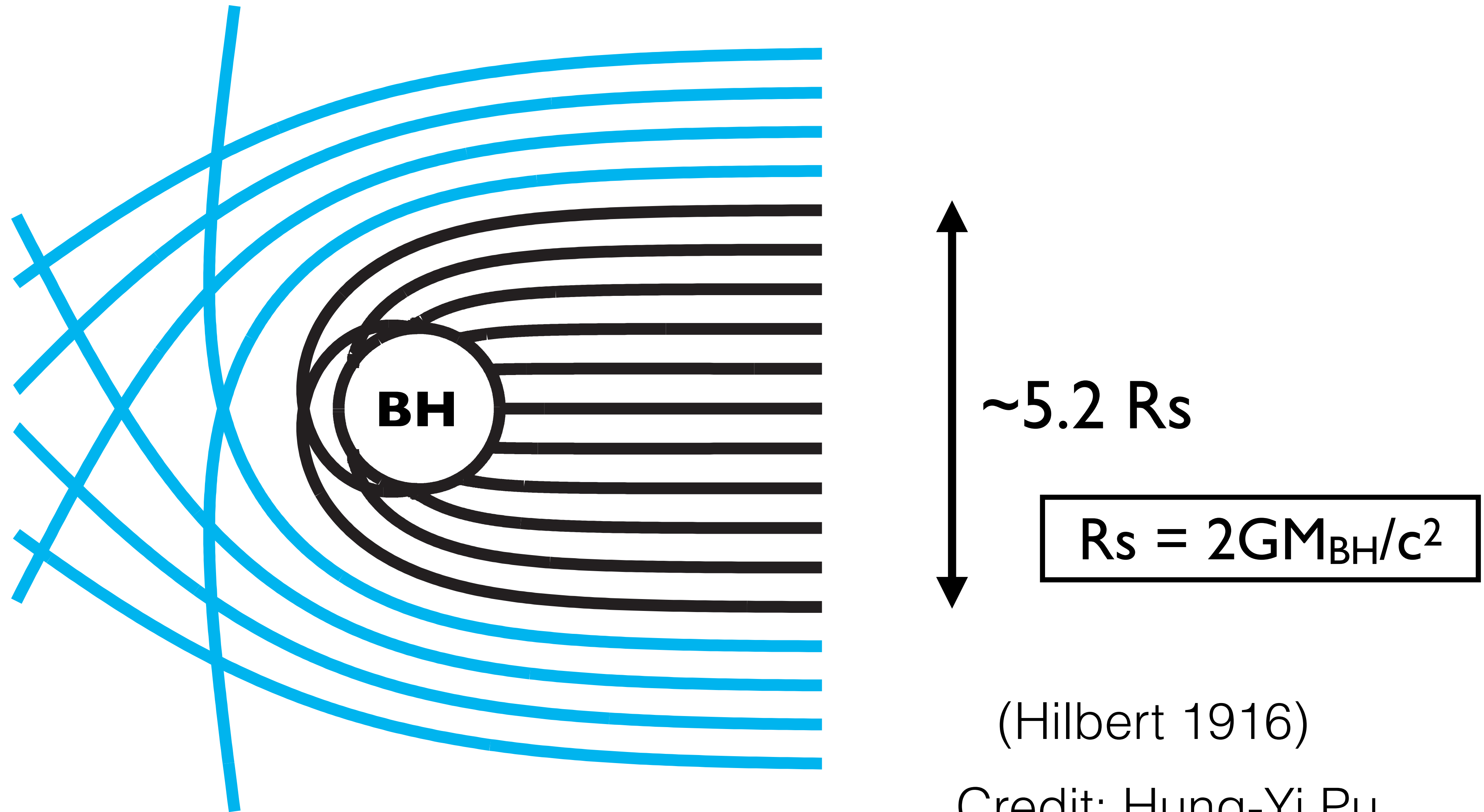
MIT Haystack Observatory



Event Horizon Telescope



The Shadow of a Black Hole



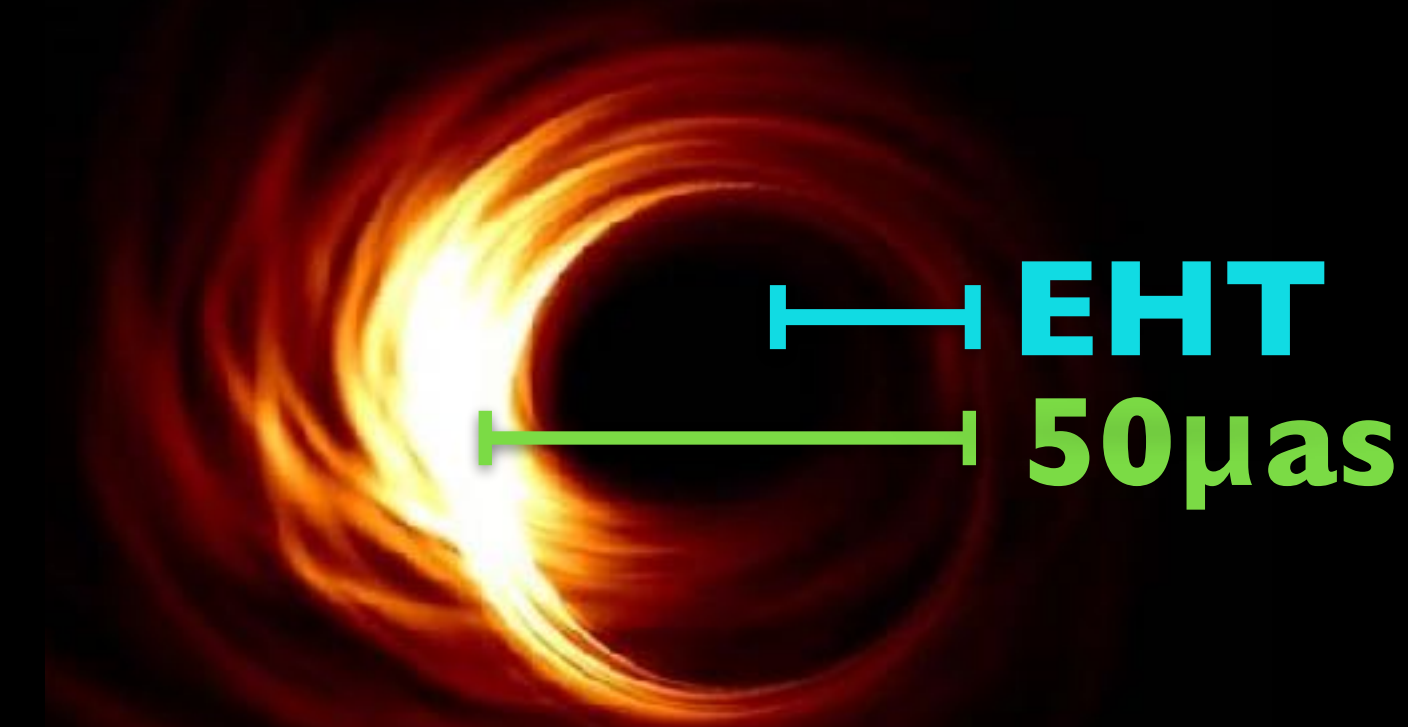
Event Horizon Telescope



Event Horizon Telescope

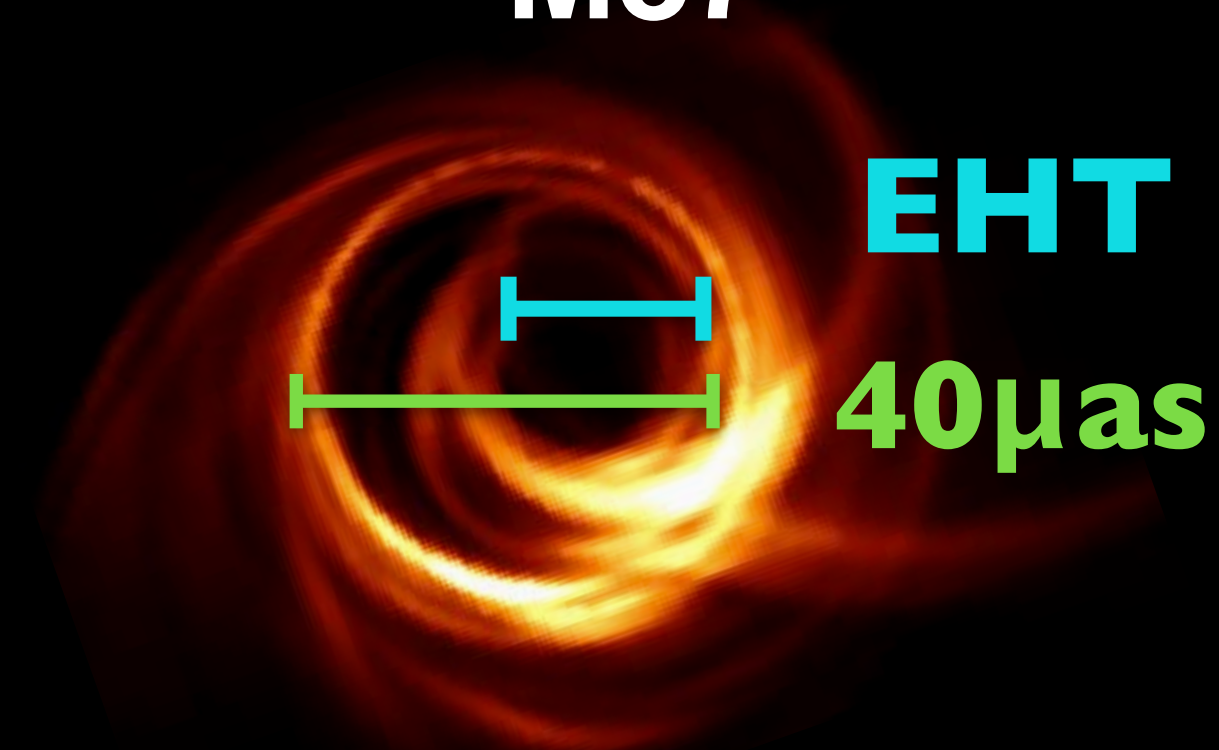


Sgr A*



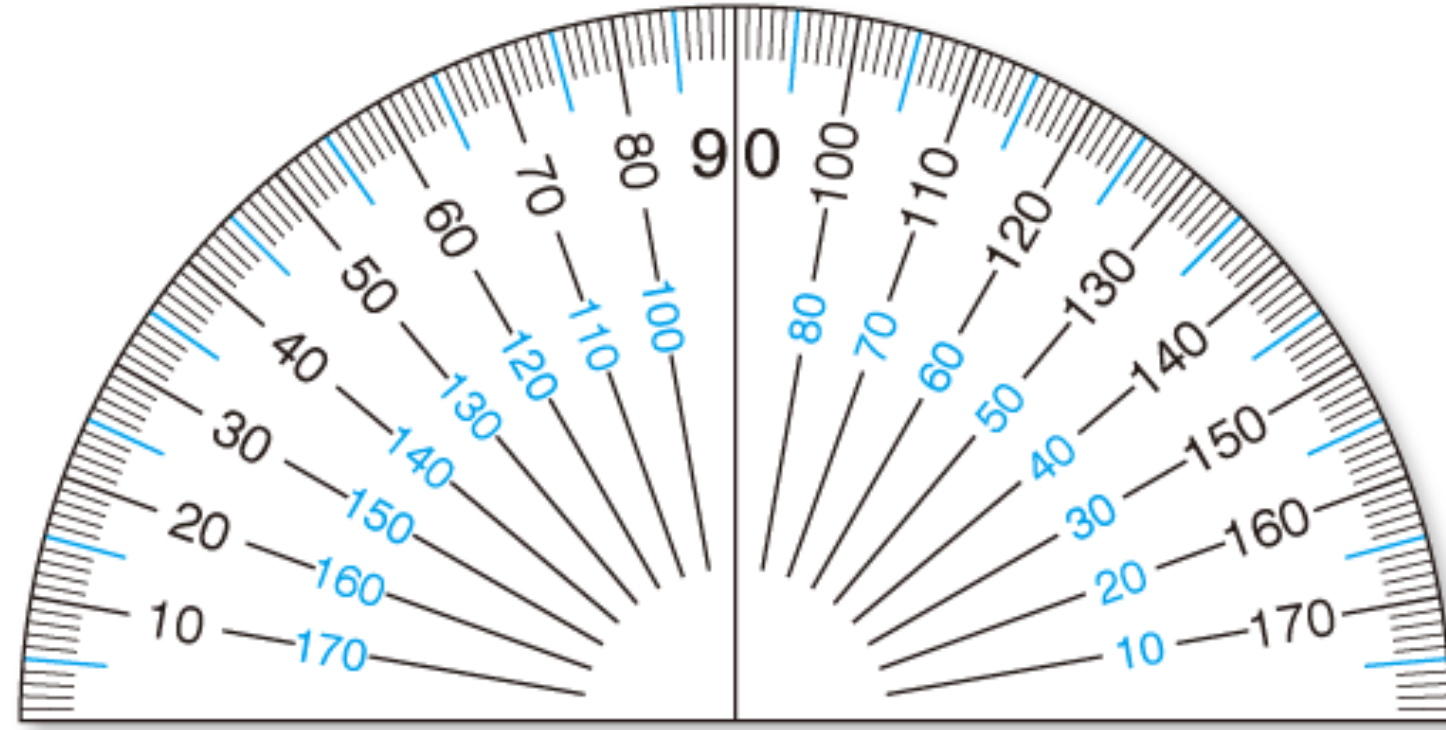
Credit: Hotaka Shiokawa

M87



Credit: Monika Moscibrodzka

Units of the Angular Size



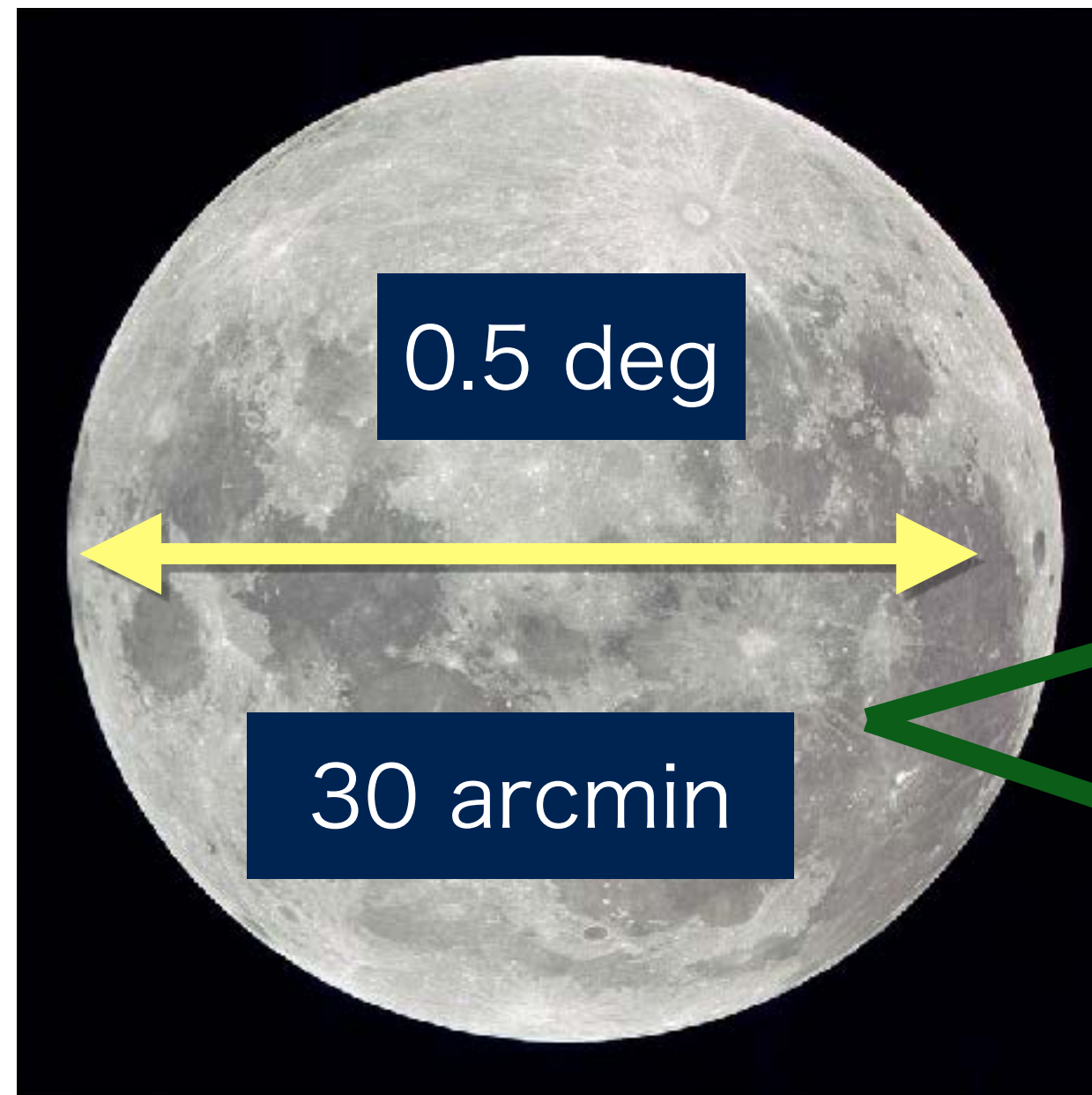
Protractor : 1 ticks = 1 degree

$\times 1/60 = 1$ arcmin

$\times 1/60 = 1$ arcsec

$\times 1/1000 = 1$ mas

$\times 1/1000 = 1$ μ as



40 - 50 μ as

Event Horizon Telescope Collaboration



>300 members, >59 institutes, >18 countries in North & South America, Europe, Asia, and Africa.



Event Horizon Telescope

Kazu Akiyama, NEROC Symposium 2020, Online, 2020/11/17 (Tue)

Meet the Telescope

SMT, Arizona



LMT, Mexico



IRAM 30m Spain



JCMT, Hawaii



APEX, Chile



Photos: ALMA, Sven Dornbusch, Junhan Kim, Helge Rottmann, David Sanchez, Daniel Michalik, Jonathan Weintroub, William Montgomerie

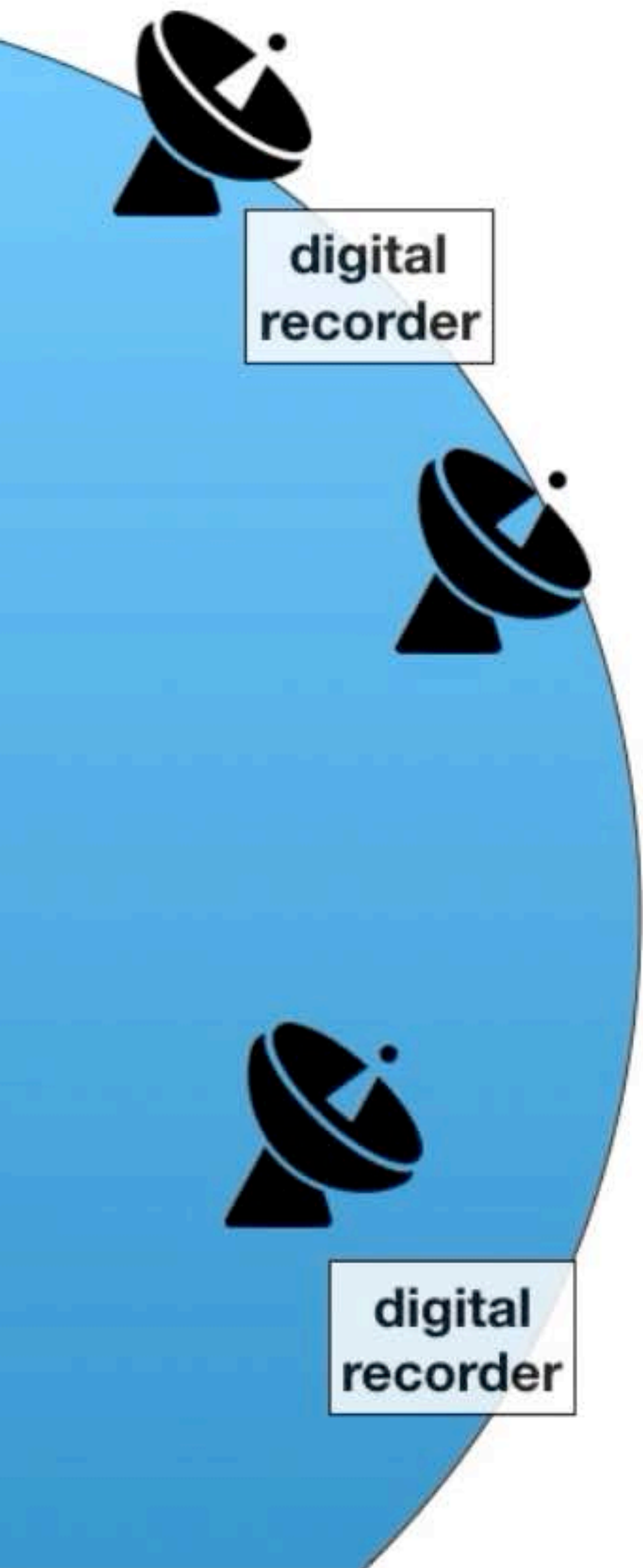
Meet the Telescope



Photos: ALMA, Sven Dornbusch, Junhan Kim, Helge Rottmann, David Sanchez, Daniel Michalik, Jonathan Weintroub, William Montgomerie

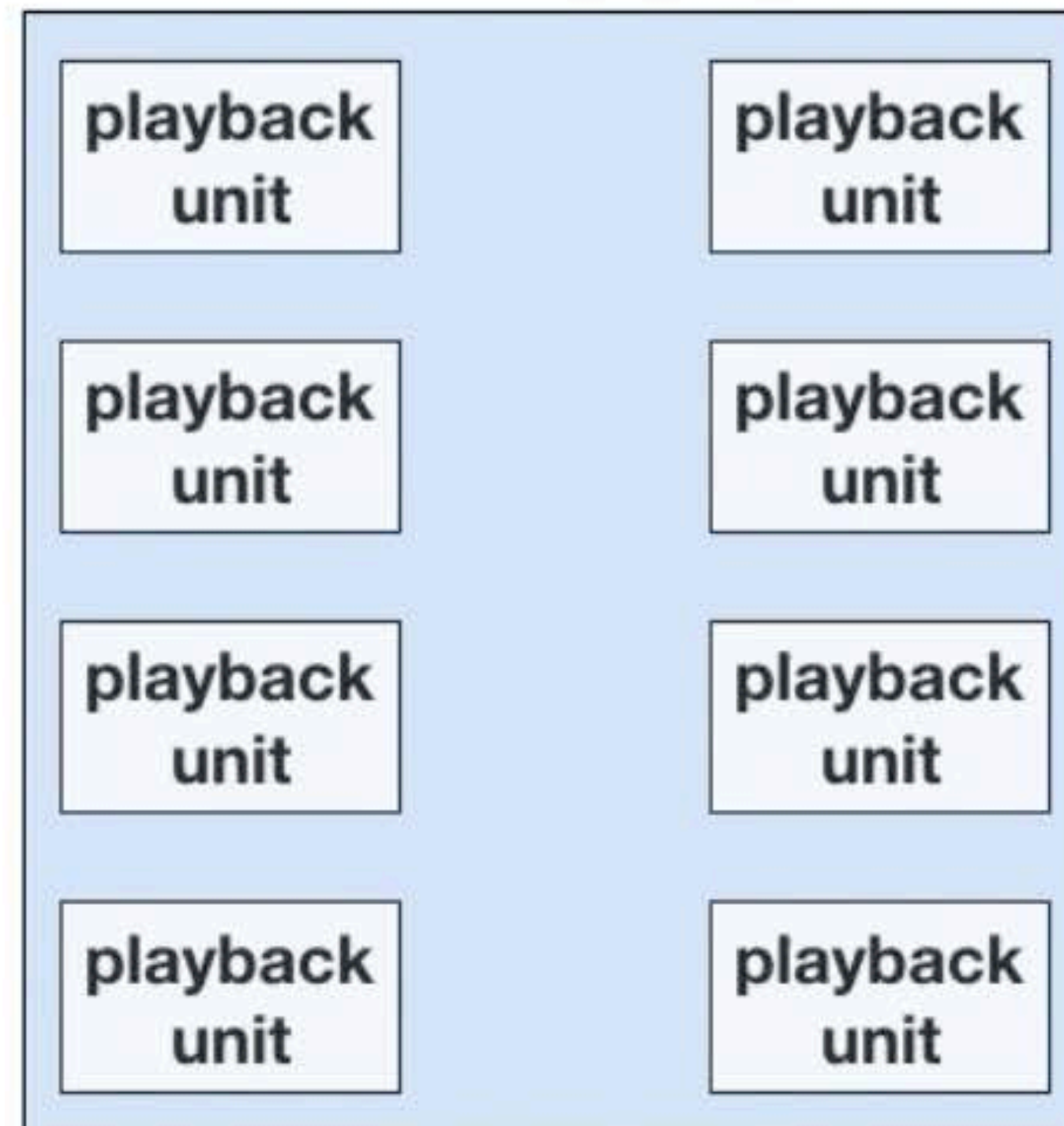


From Observations to Images

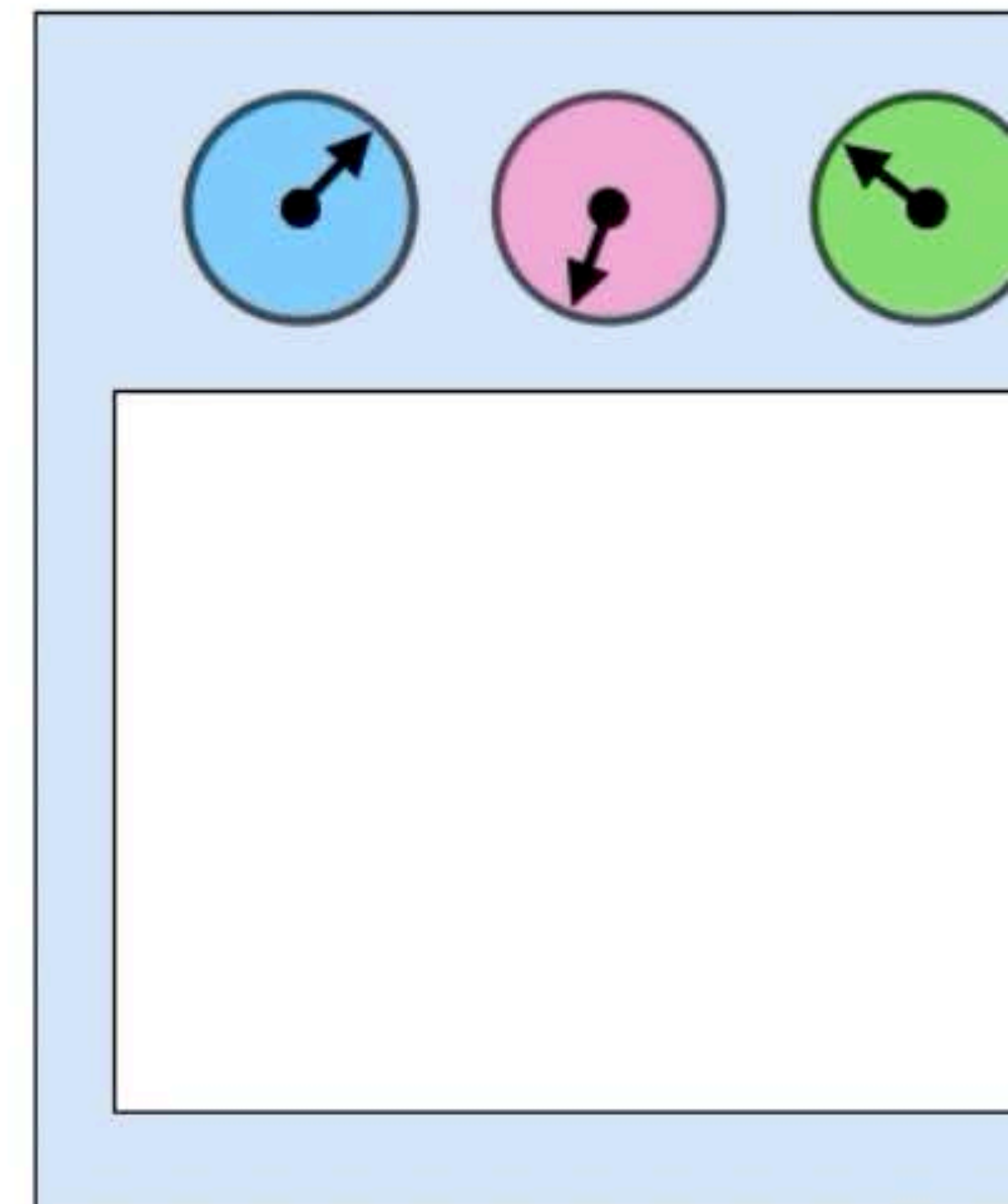


digital recorder

EHT correlator

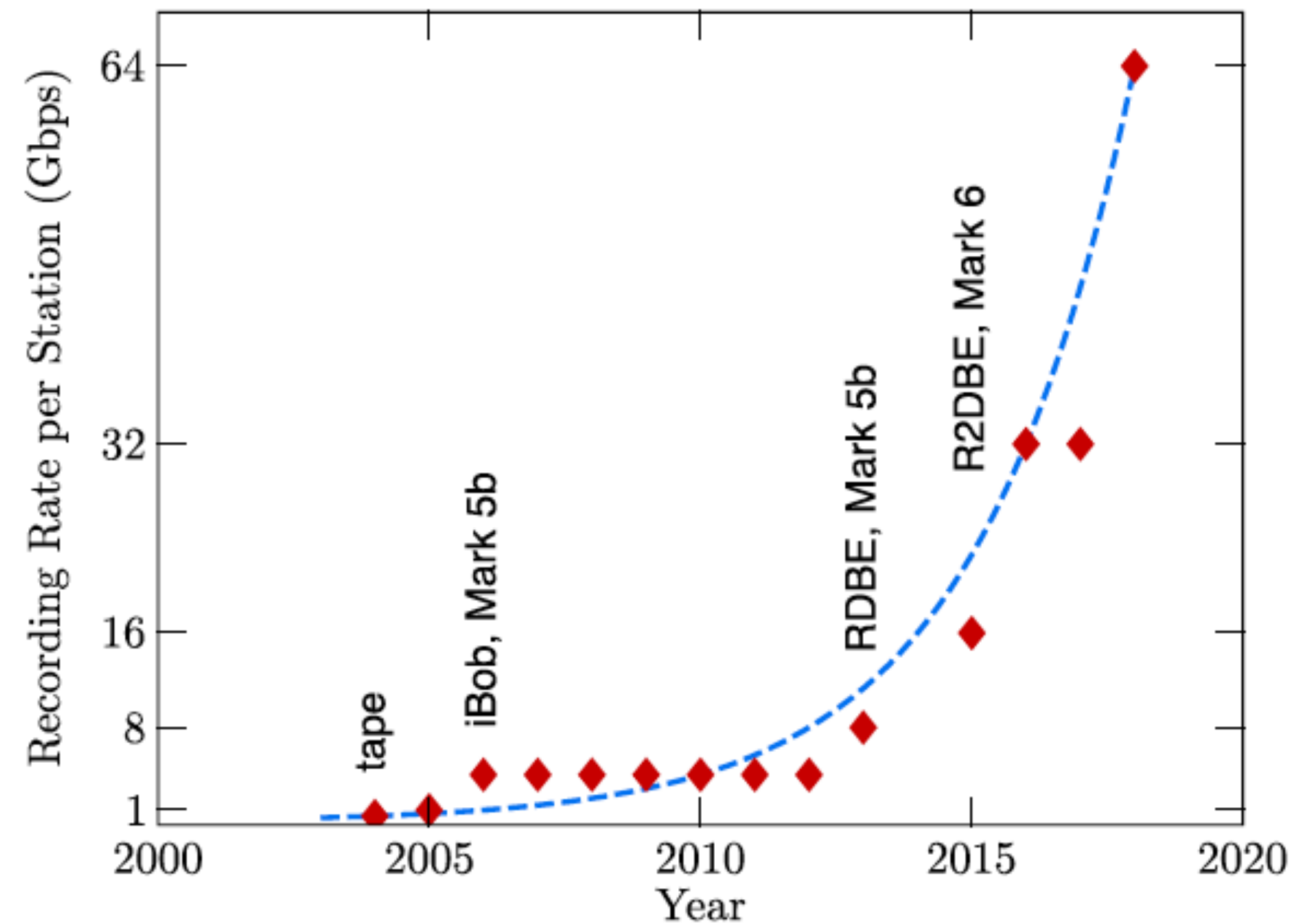


Calibration



EHT Hardware

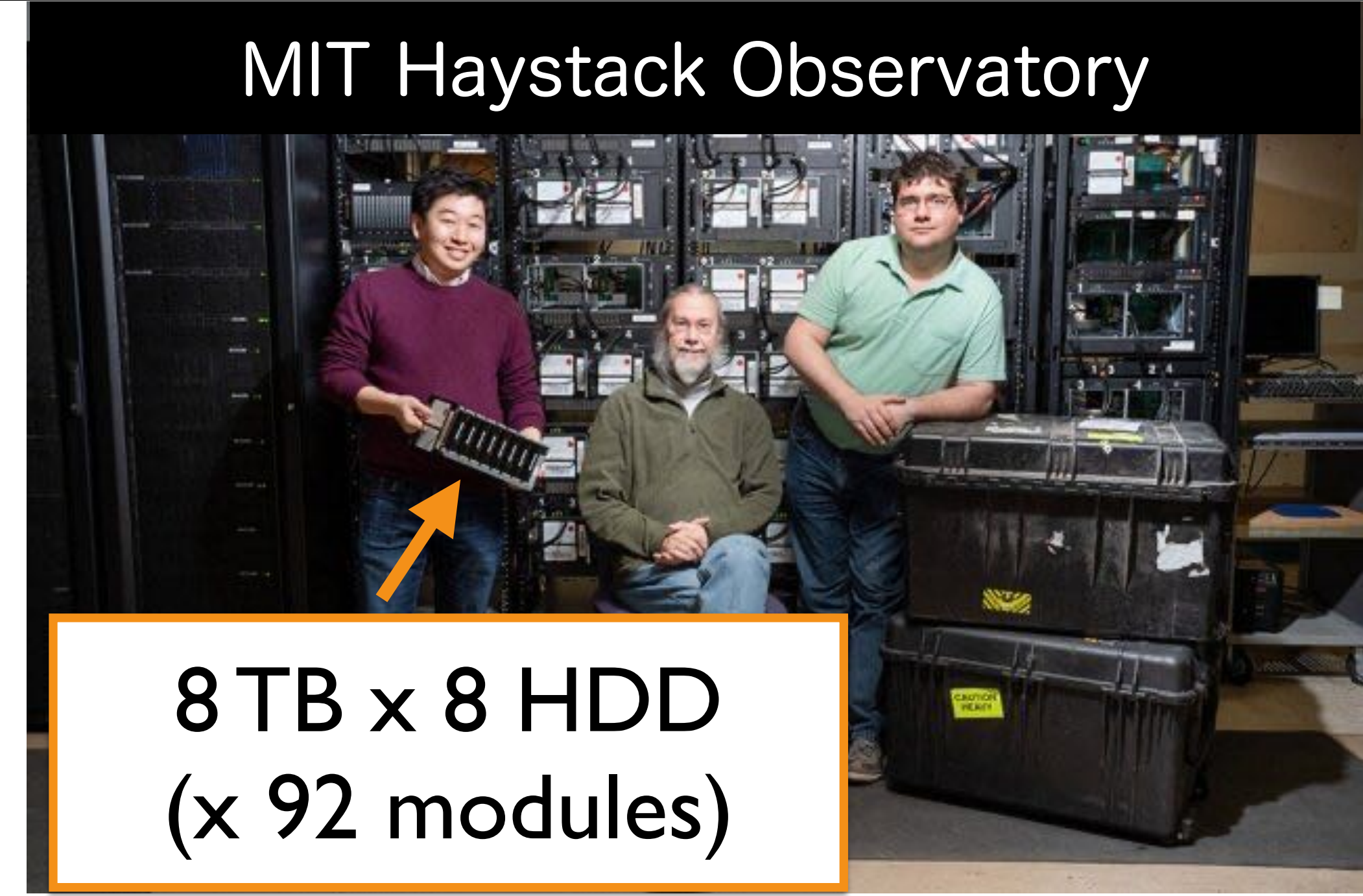
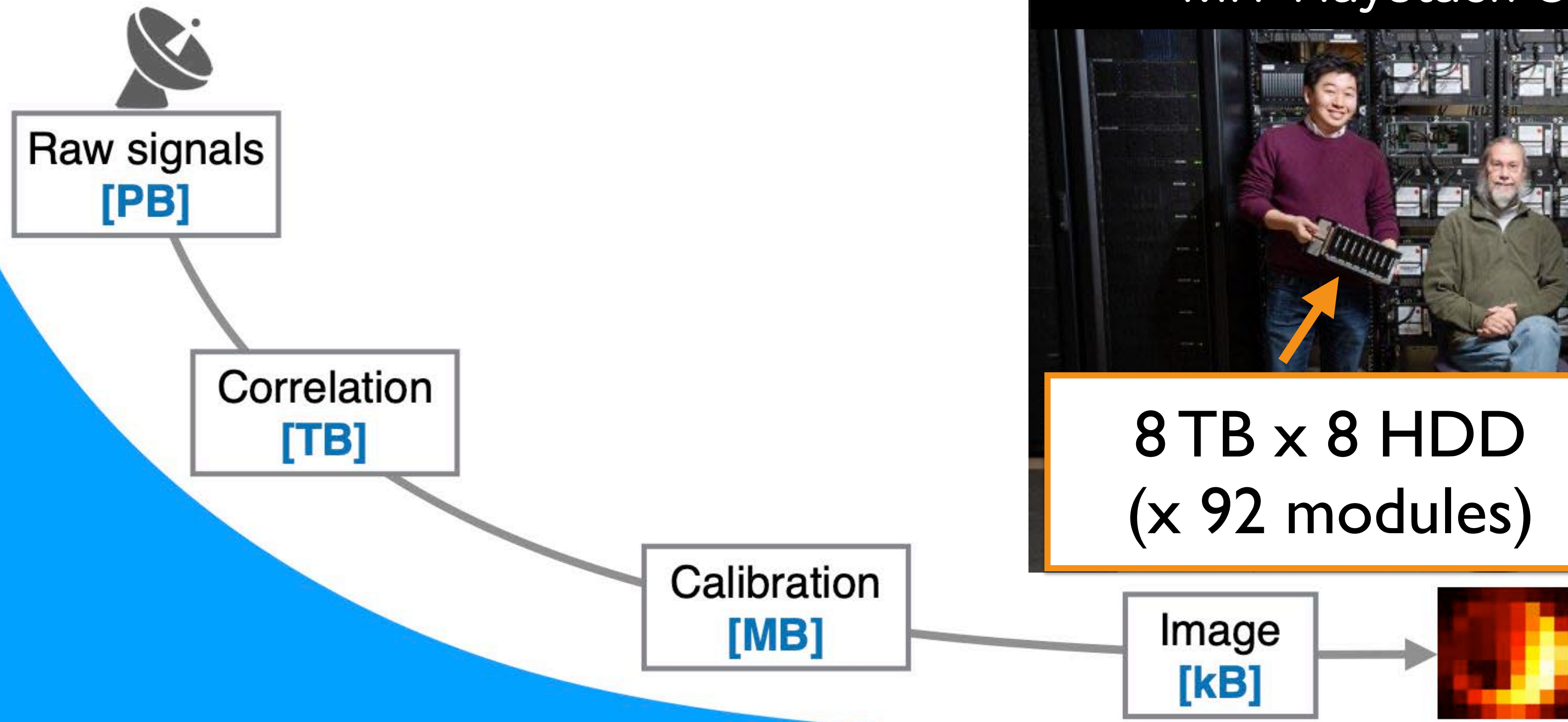
EHT Backend
(R2DBE, Mark 6)



Recording Rate:

- VLBA, GMVA: 2-4 Gbps
- EHT: 32 Gbps (2017), 64 Gbps (2018-)

From Observations to Images

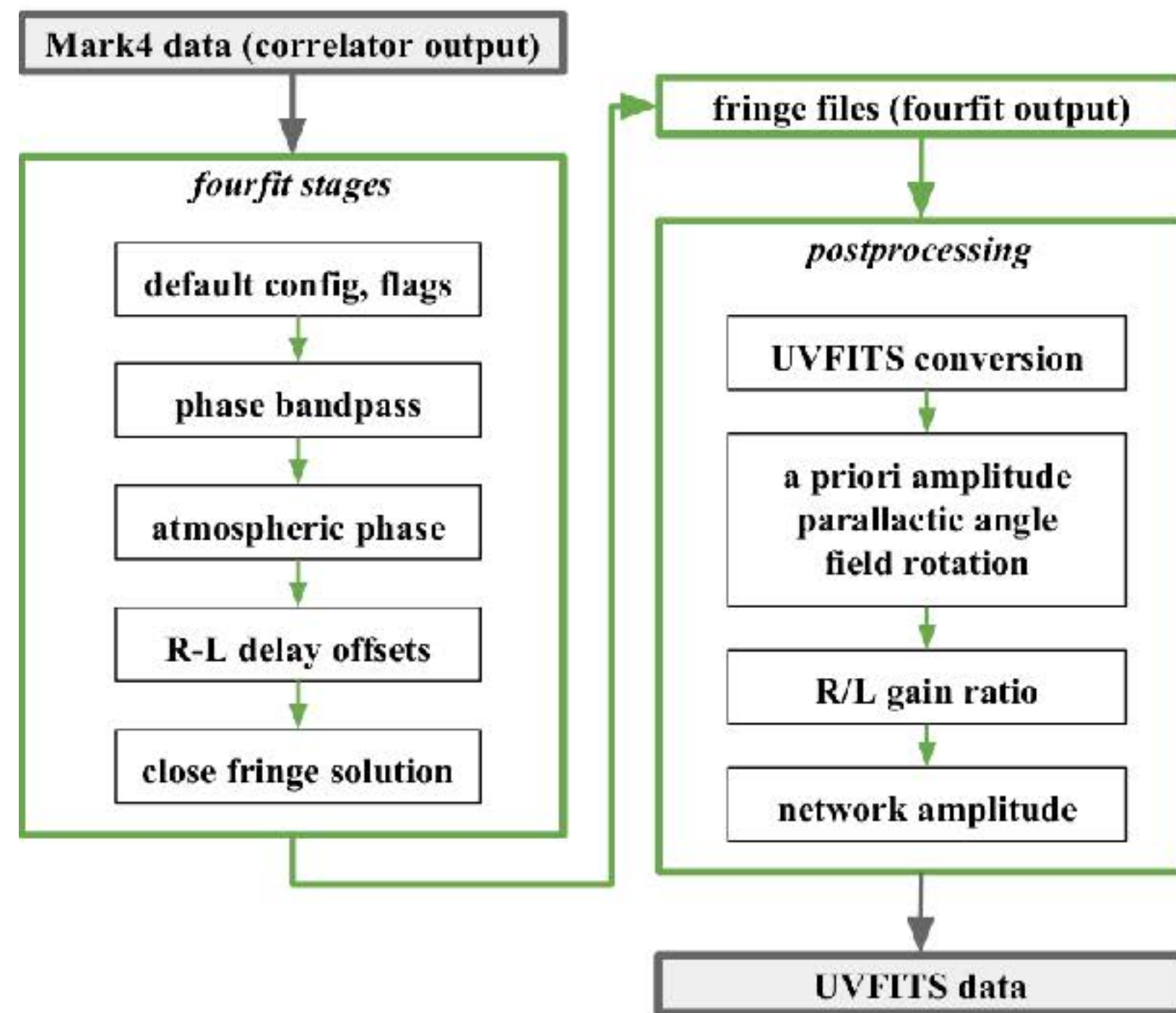


Credit: Bryce Vickmark

Credit: Lindy Blackburn

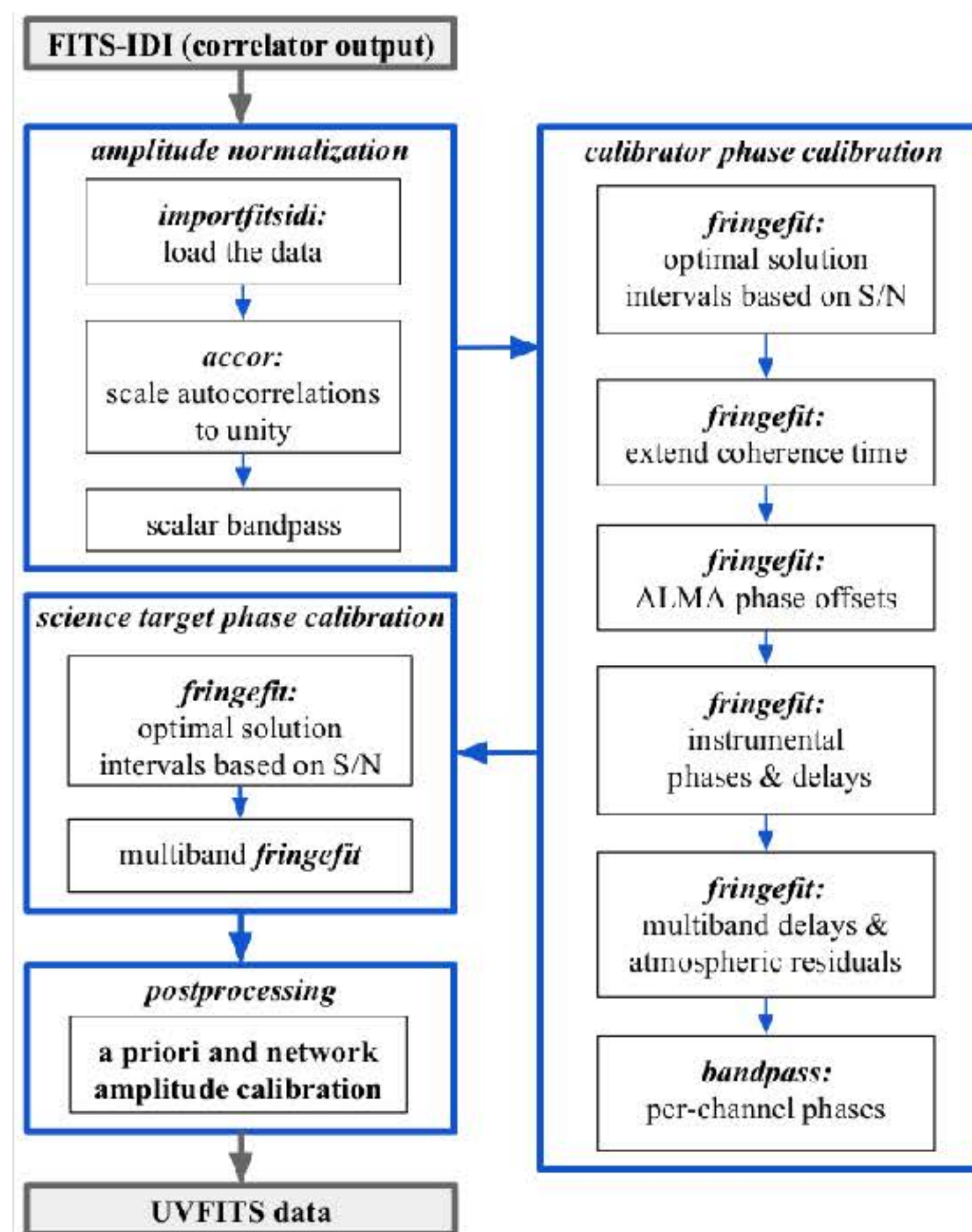
Data Calibration

HOPS Pipeline (EHT-HOPS)



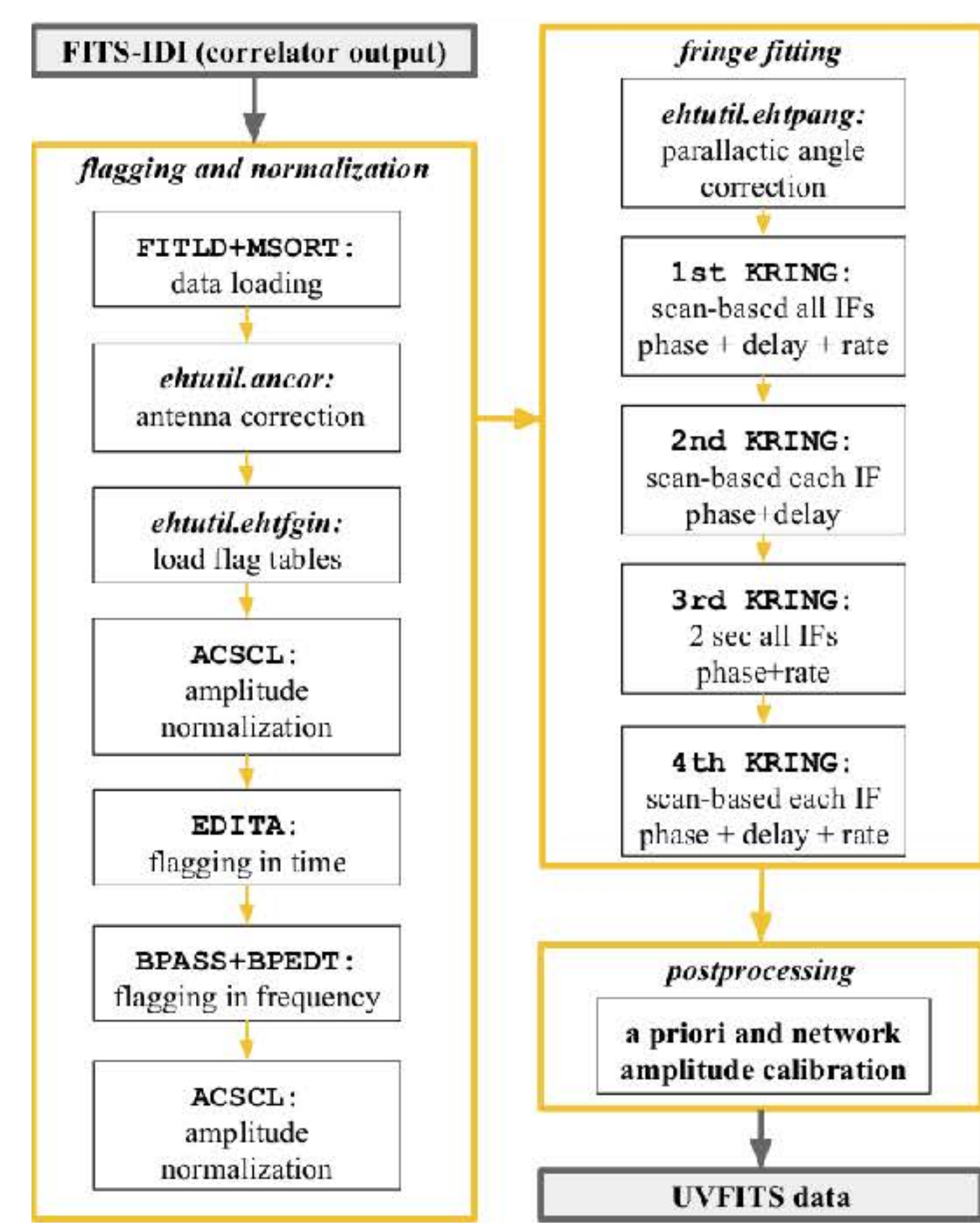
Blackburn et al. 2019

CASA Pipeline (rPICARD)



Janssen et al. 2019

EHT AIPS Pipeline



EHT Collaboration 2019, ApJL, 875, L3

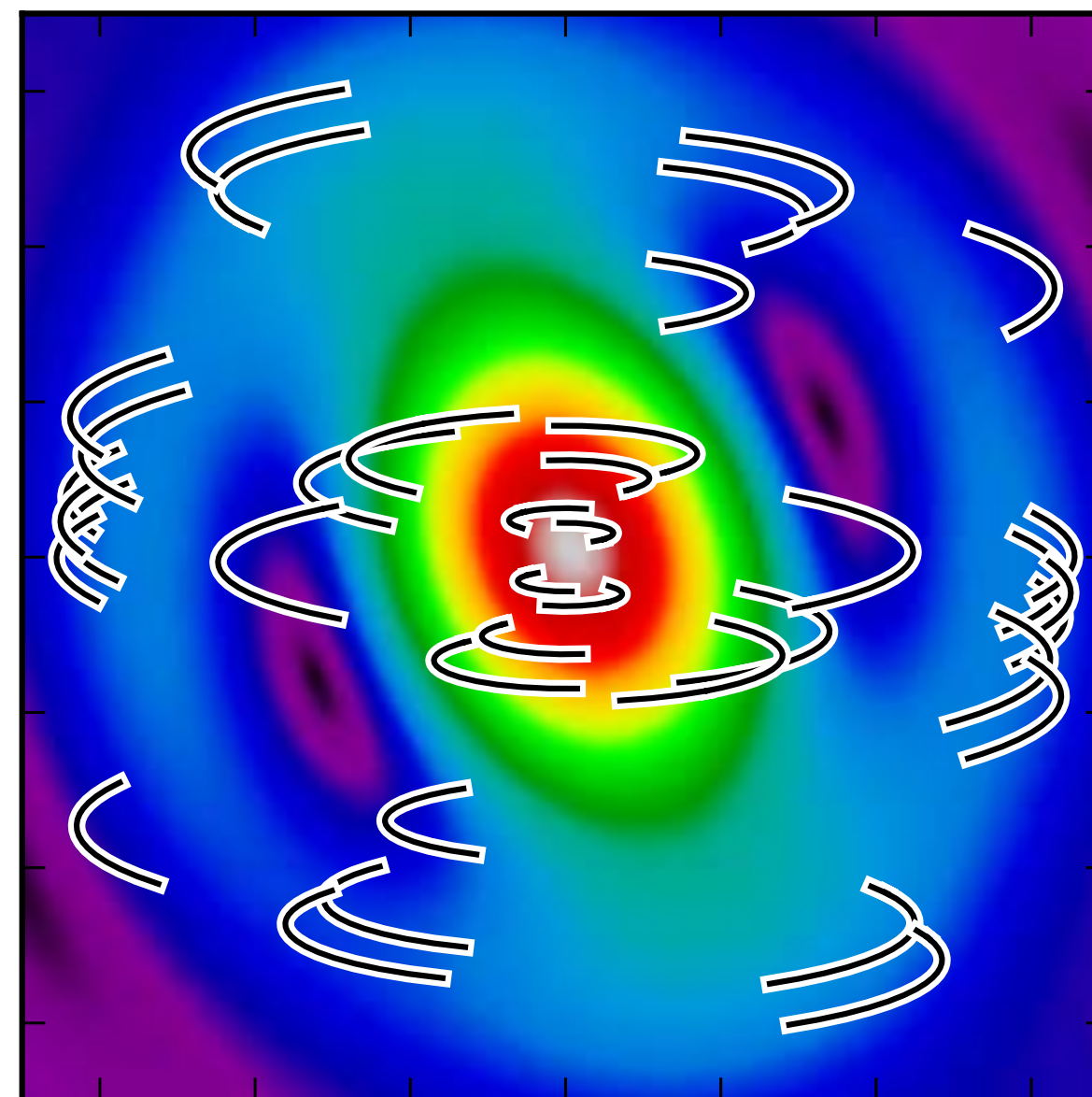


Interferometric Imaging

Image

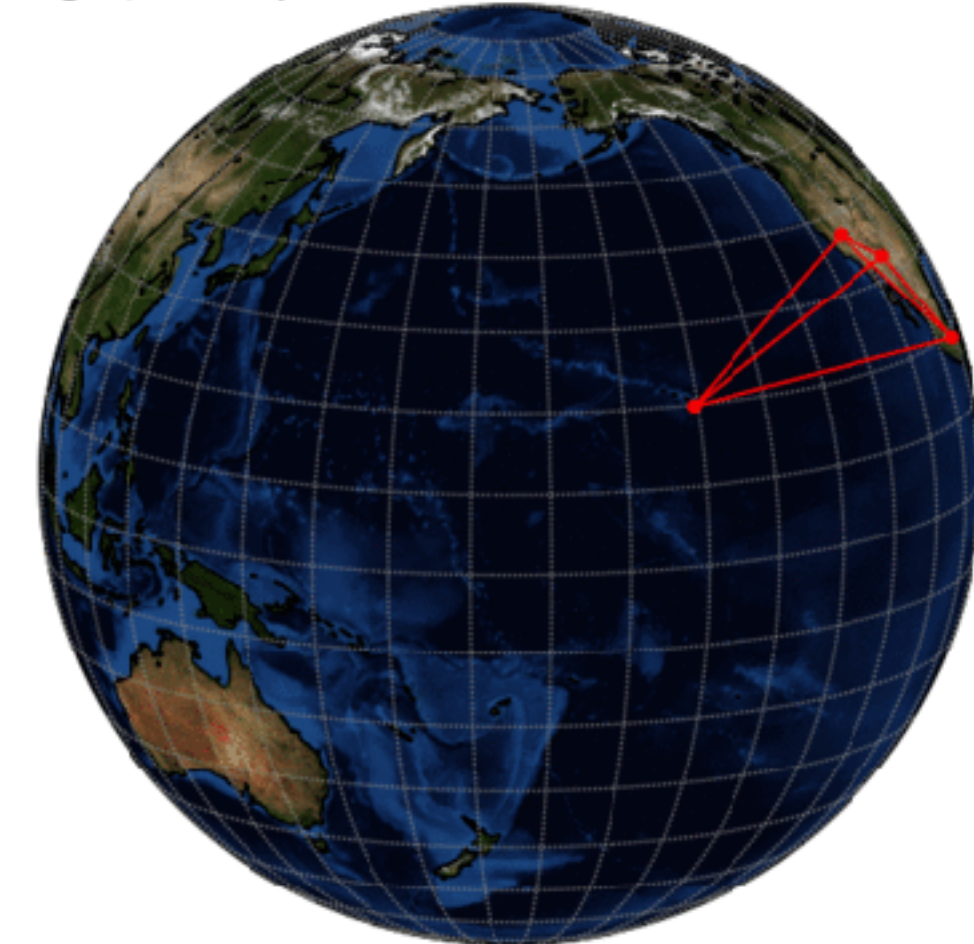


**Fourier Domain
(Visibility)**



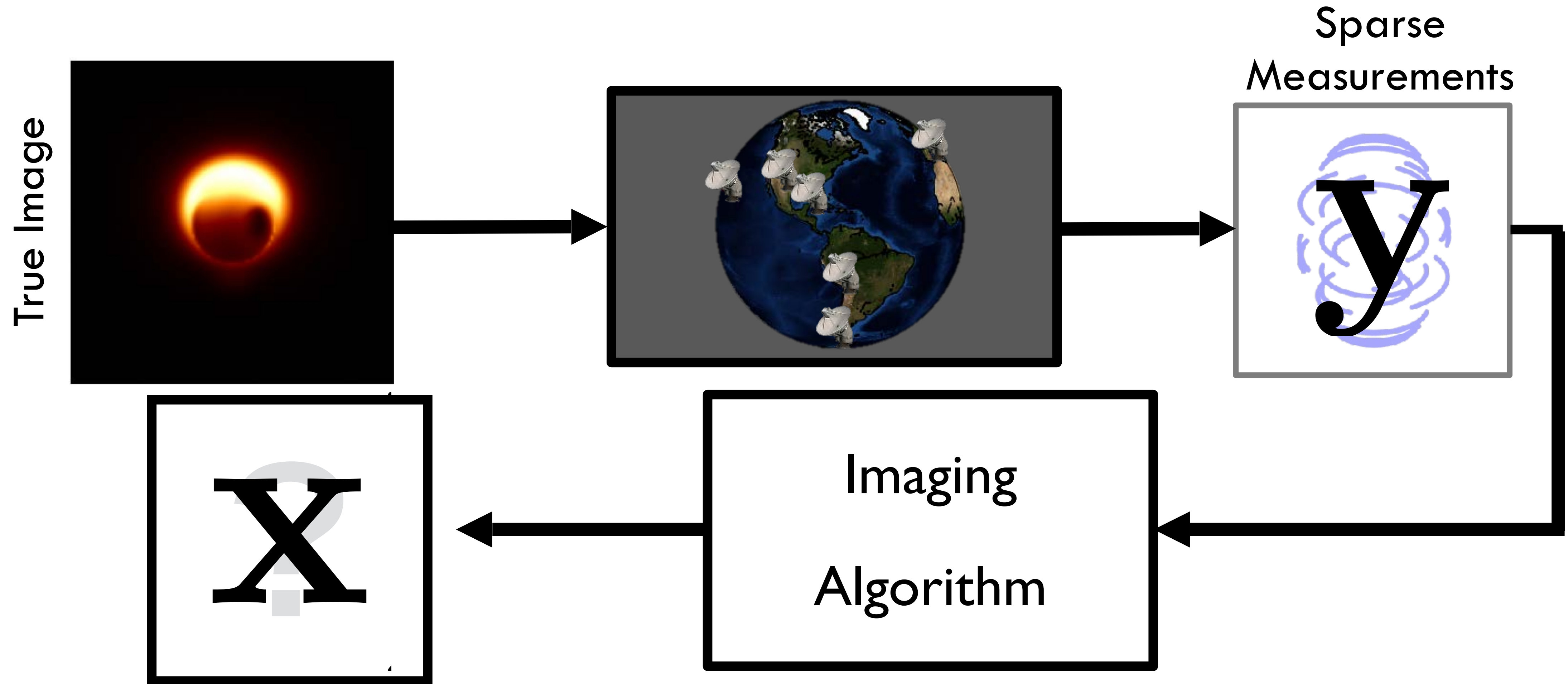
**Sampling Process
(Projected Baseline = Spatial Frequency)**

Orthographic Map Centered on Lon=180, Lat=12.391123



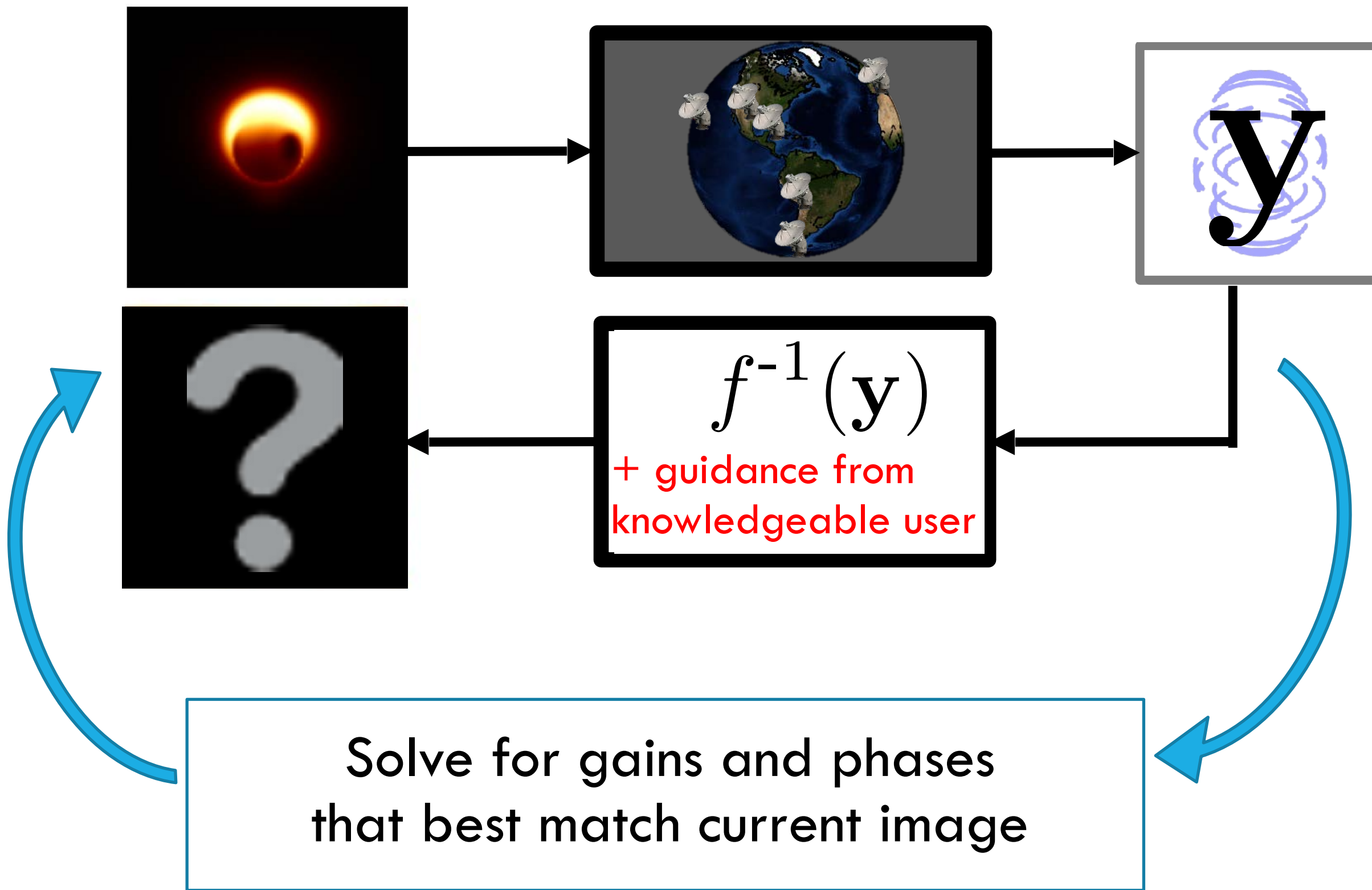
(Images: adapted from **Akiyama** et al. 2015, ApJ ; Movie: Laura Vertatschitsch)

Interferometric Imaging



Two Classes of Imaging Algorithms

Credit: Katie Bouman



Traditional

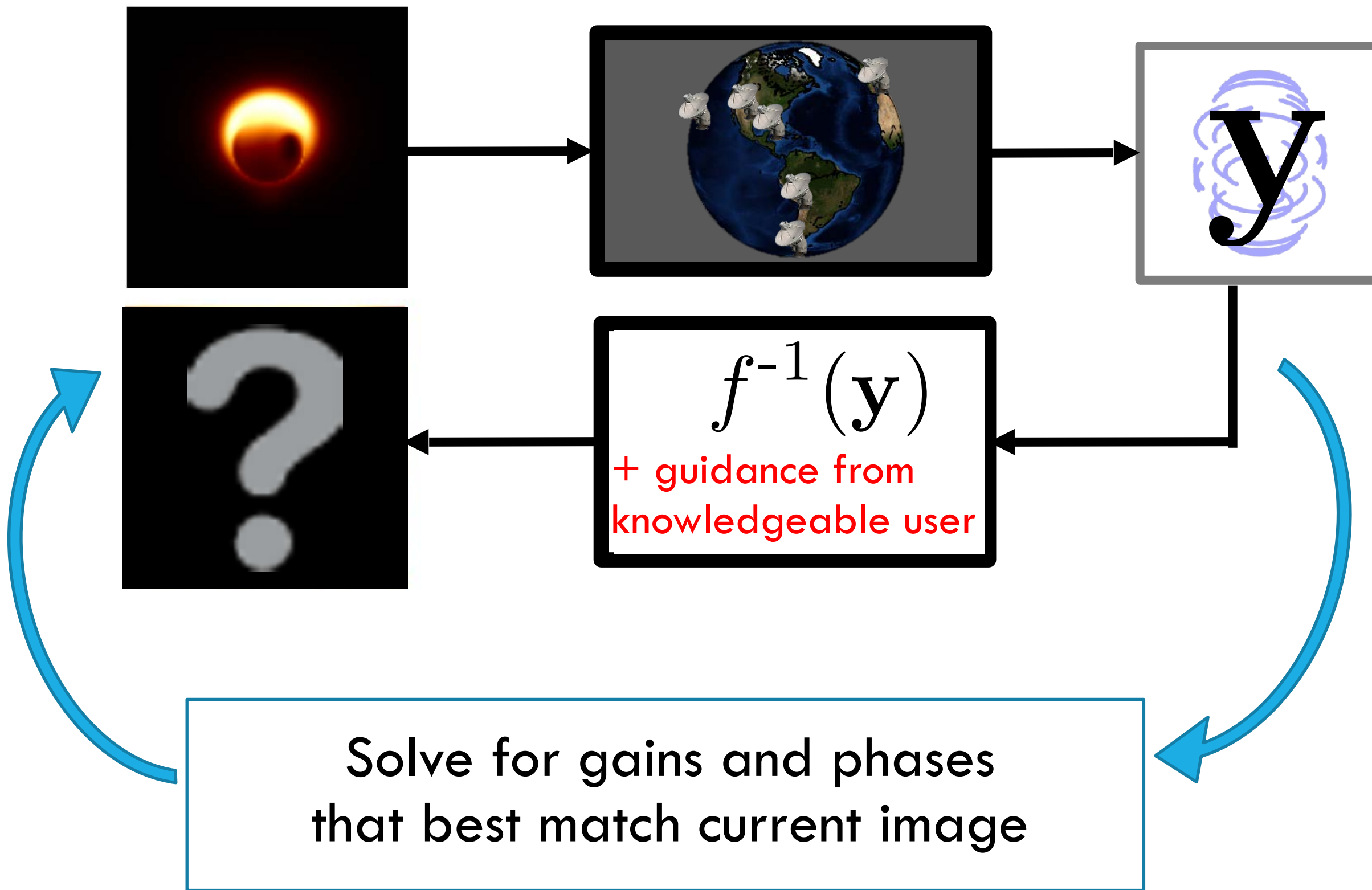
Inverse Modeling
(CLEAN + Self-Calibration)



Event Horizon Telescope

Two Classes of Imaging Algorithms

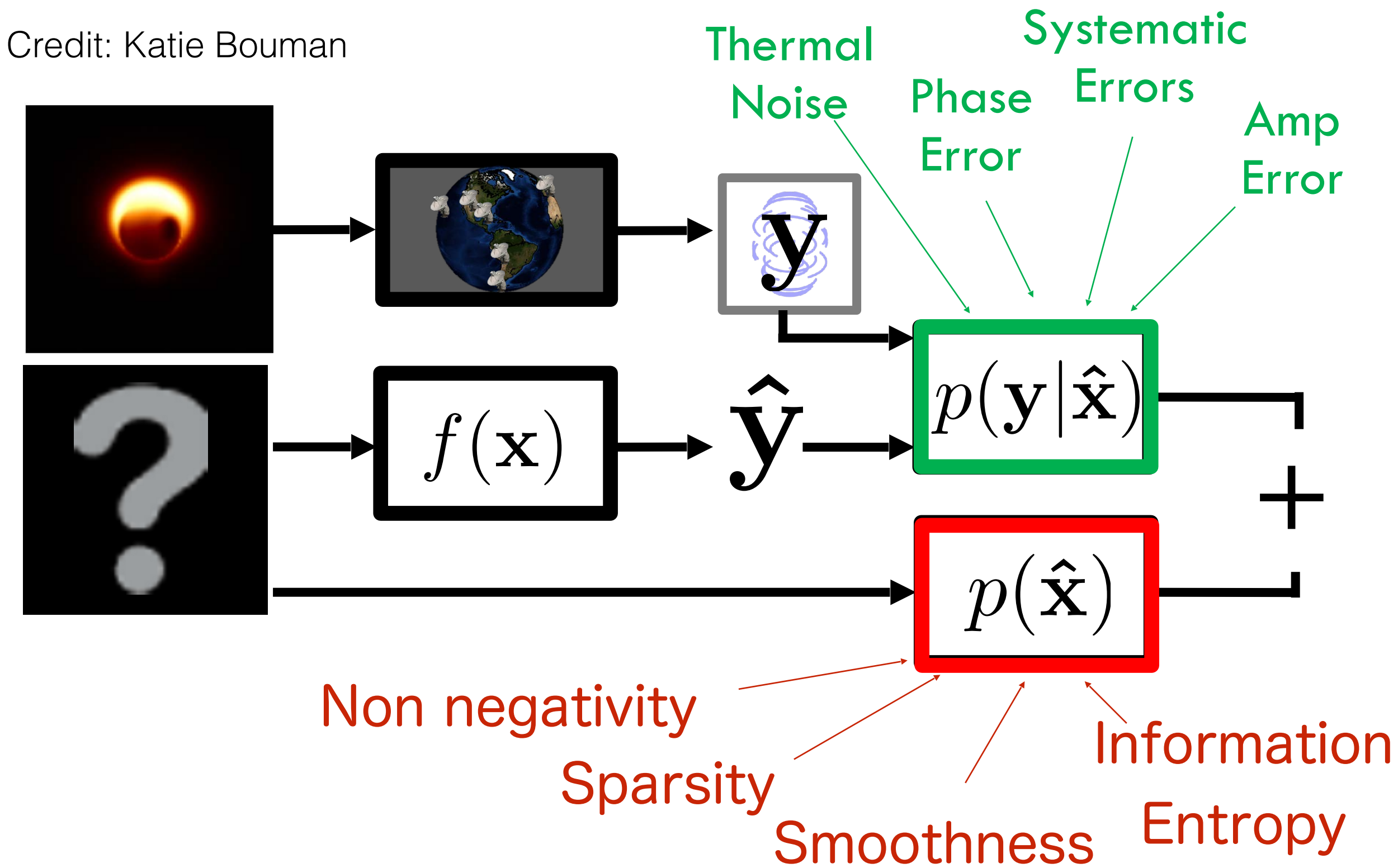
Credit: Katie Bouman



Traditional

**Inverse Modeling
(CLEAN + Self-Calibration)**

Credit: Katie Bouman



$$\hat{\mathbf{x}}_{\text{MAP}} = \operatorname{argmax}_{\mathbf{x}} [\log p(\mathbf{y}|\mathbf{x}) + \log p(\mathbf{x})]$$

**Forward Modeling
(Bayesian Inspired Optimization)**



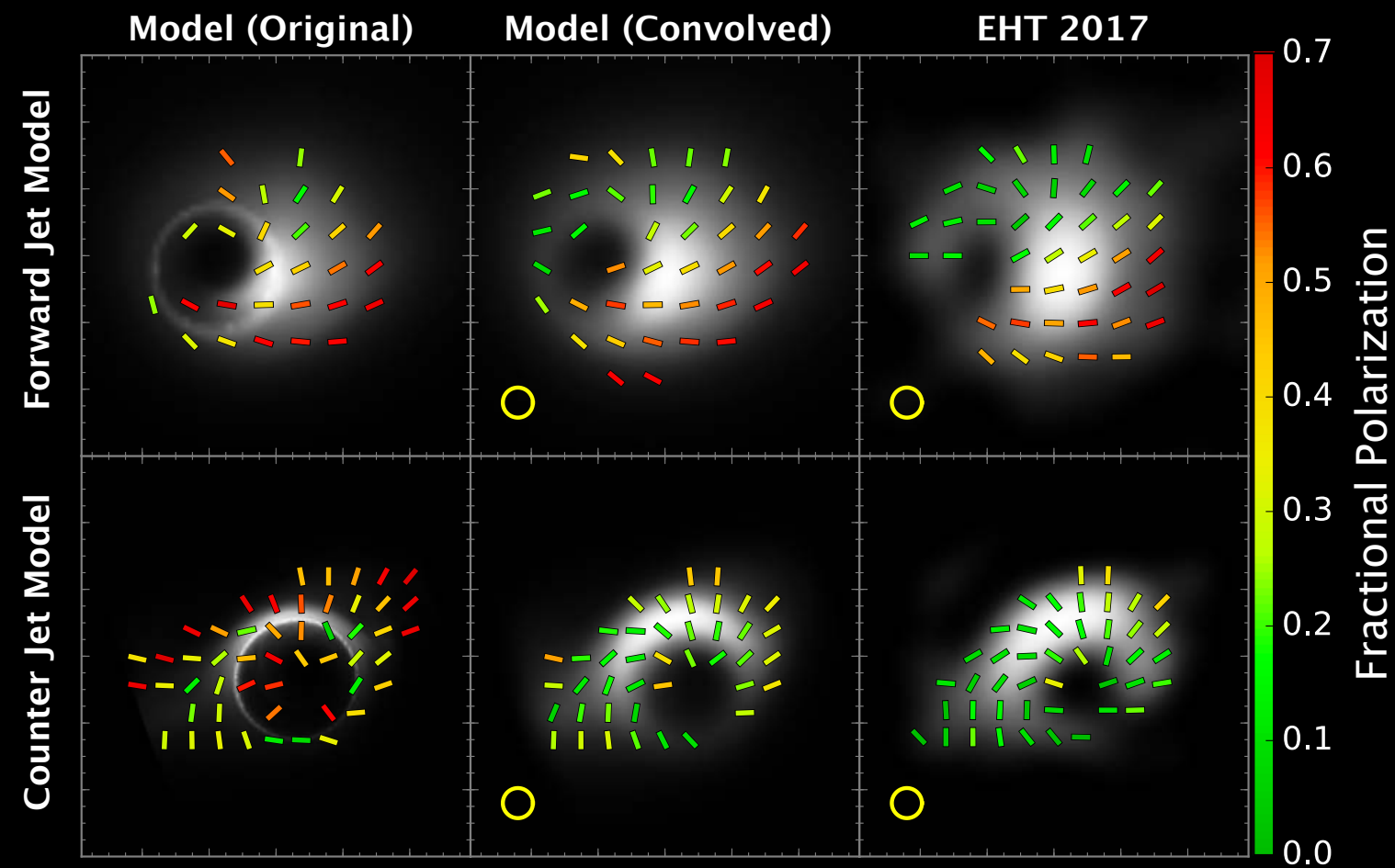
Event Horizon Telescope

New Imaging Methods

Sparse Modeling

Akiyama et al. 2017a, 2017b

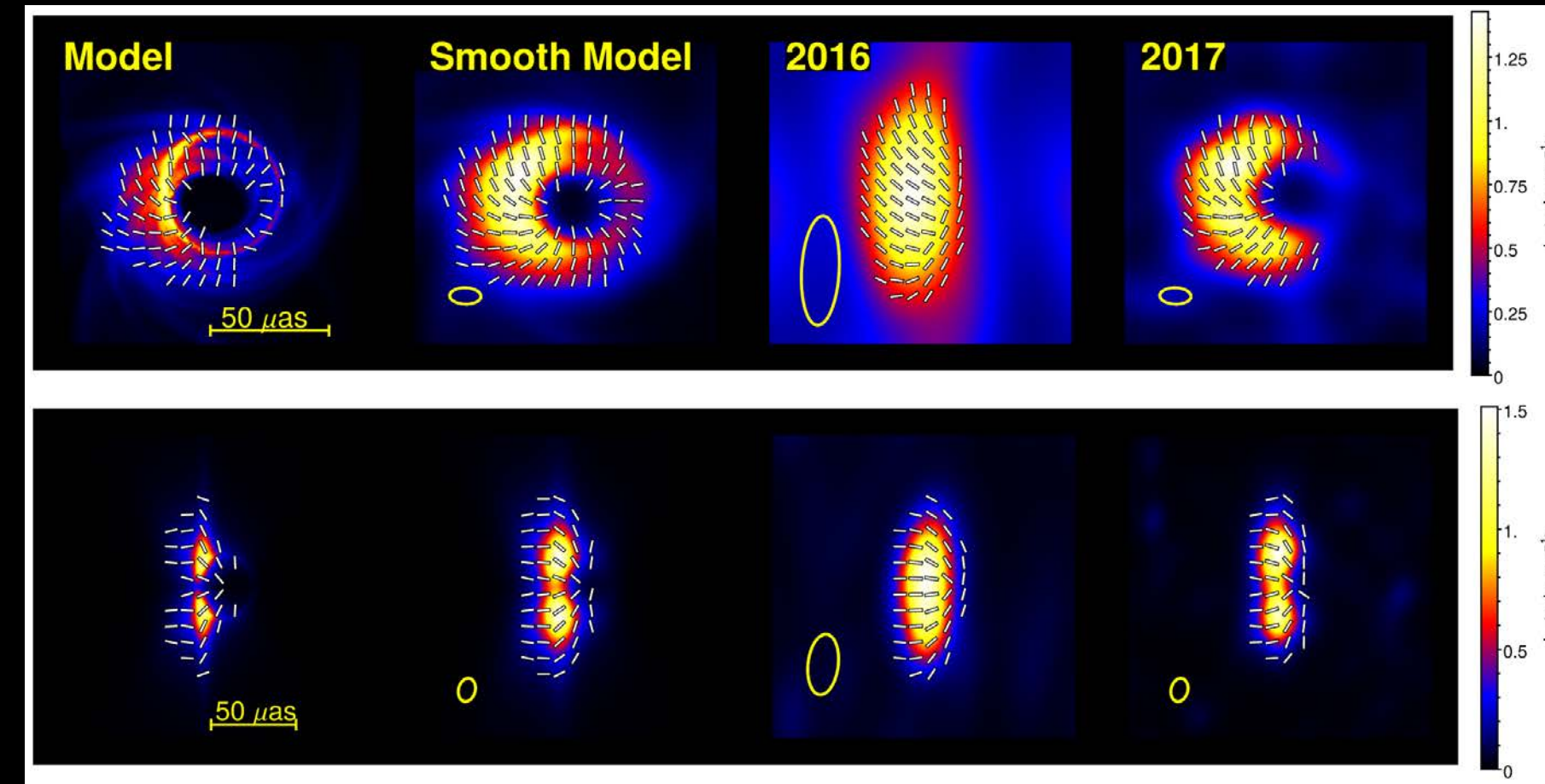
Ikeda et al. 2016, Honma et al. 2014



Maximum Entropy Method (MEM)

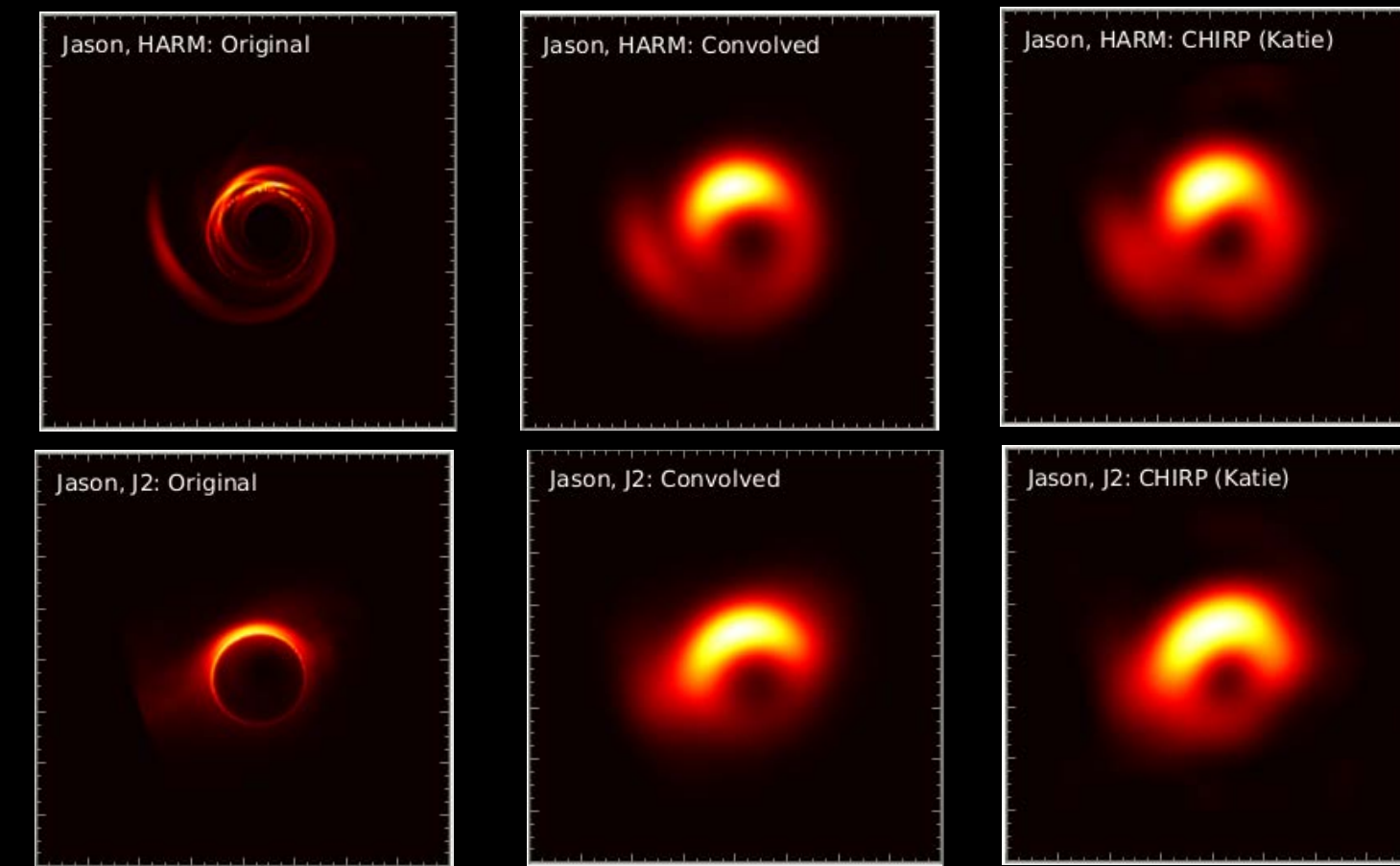
Chael et al. 2016, Fish et al. 2014,

Lu et al. 2014, 2016



CHIRP (Machine-learning)

Bouman et al. 2016



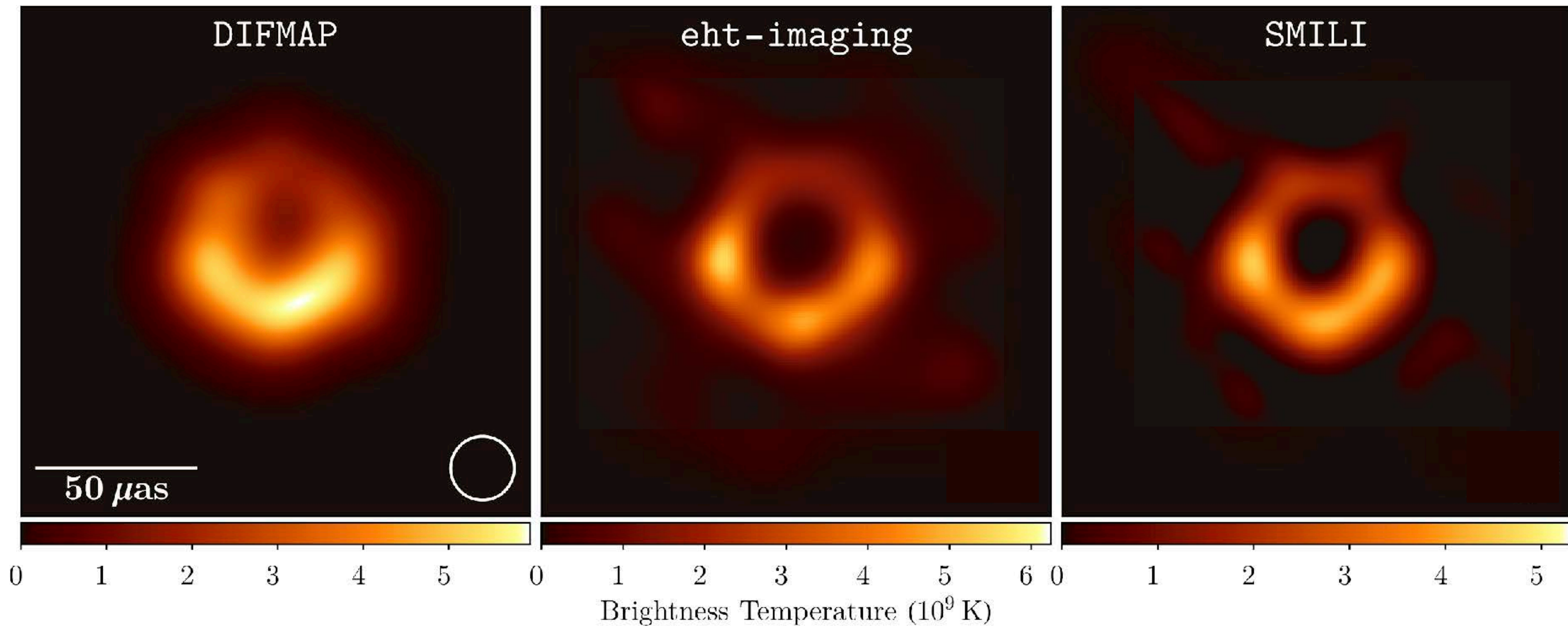
Two Imaging Libraries

eht-imaging (Chael+2016,2018) : <https://github.com/achael/eht-imaging>

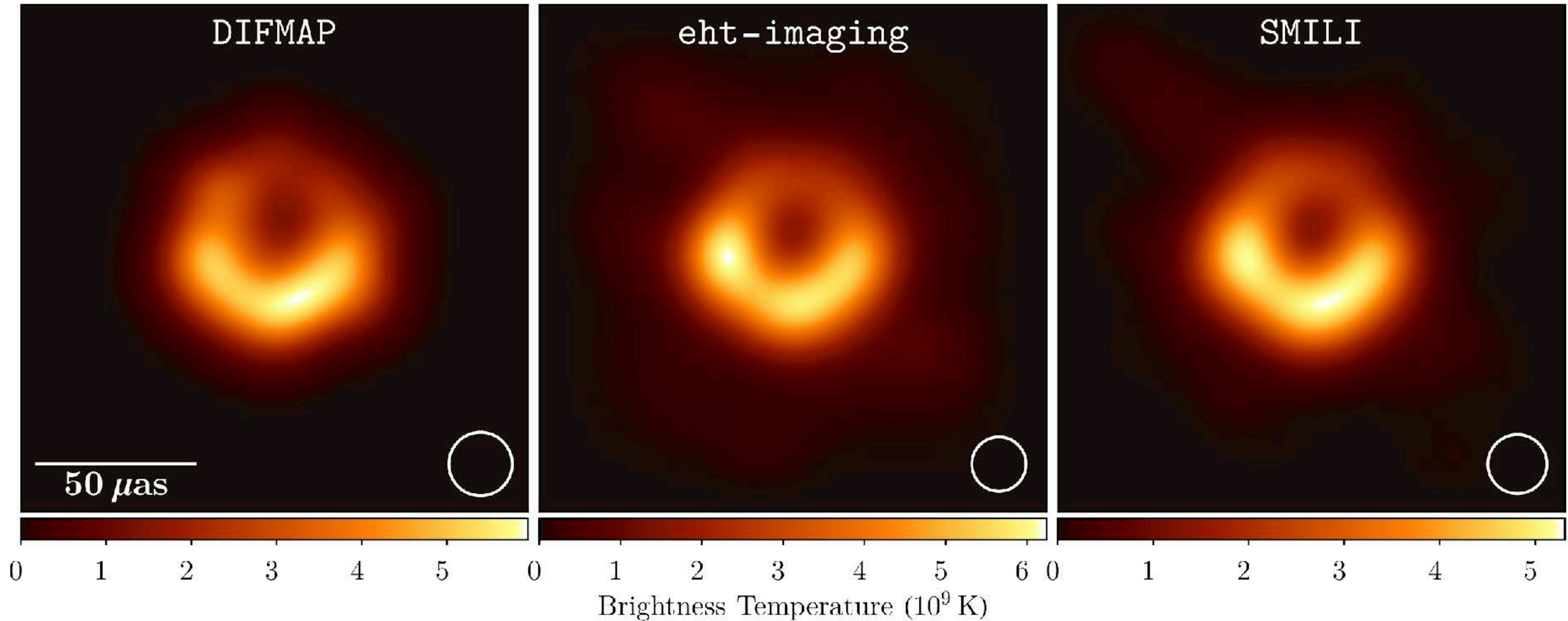
SMILI (Akiyama+2017a,b) : <https://github.com/astrosmili/smili>



Fiducial Images on Apr 11



Fiducial Images on Apr 11



✓ Stellar Mass: $6.2 \times 10^9 M_{\text{sun}}$
(Gebhardt et al. 2011)

Gas Mass: $3.5 \times 10^9 M_{\text{sun}}$
(Walsh et al. 2013)



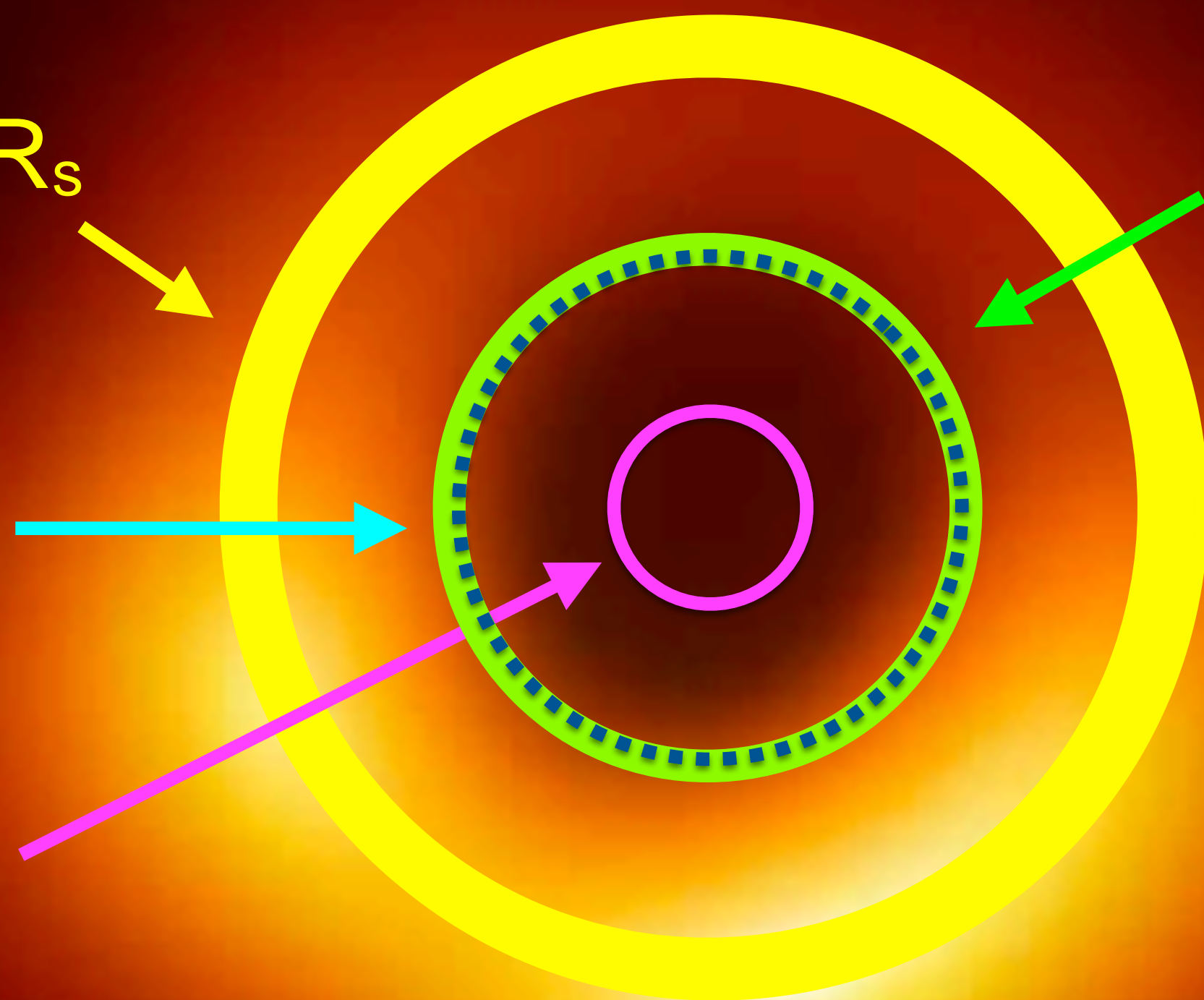
✓ Black Hole: $4.84-5.2 R_s$

Black Hole: $4.84-5.2 R_s$



✗ A worm Hole: $\sim 2.7 R_s$

✗ Naked Singularity: $1 R_s$
(extremely spinning)



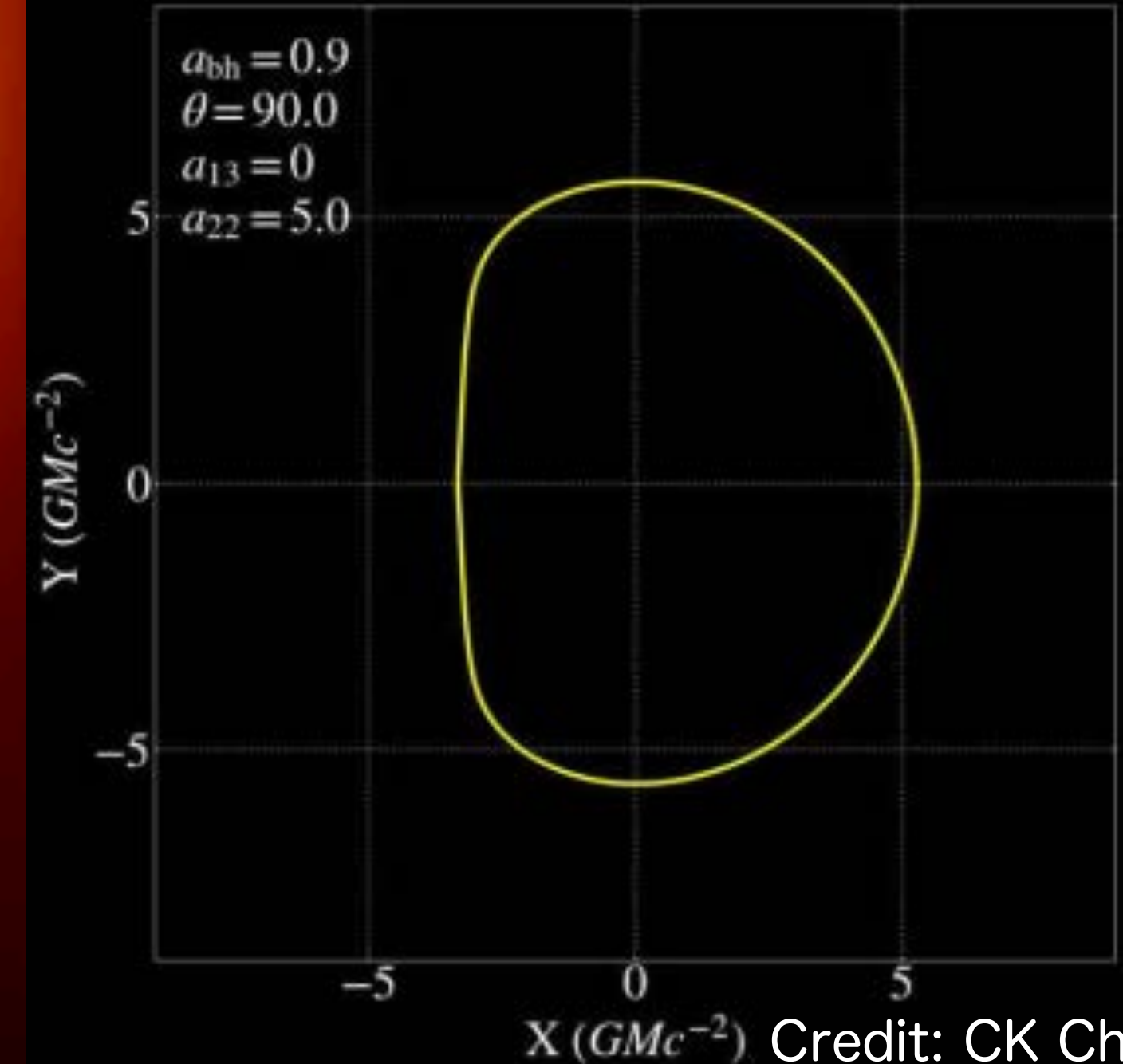
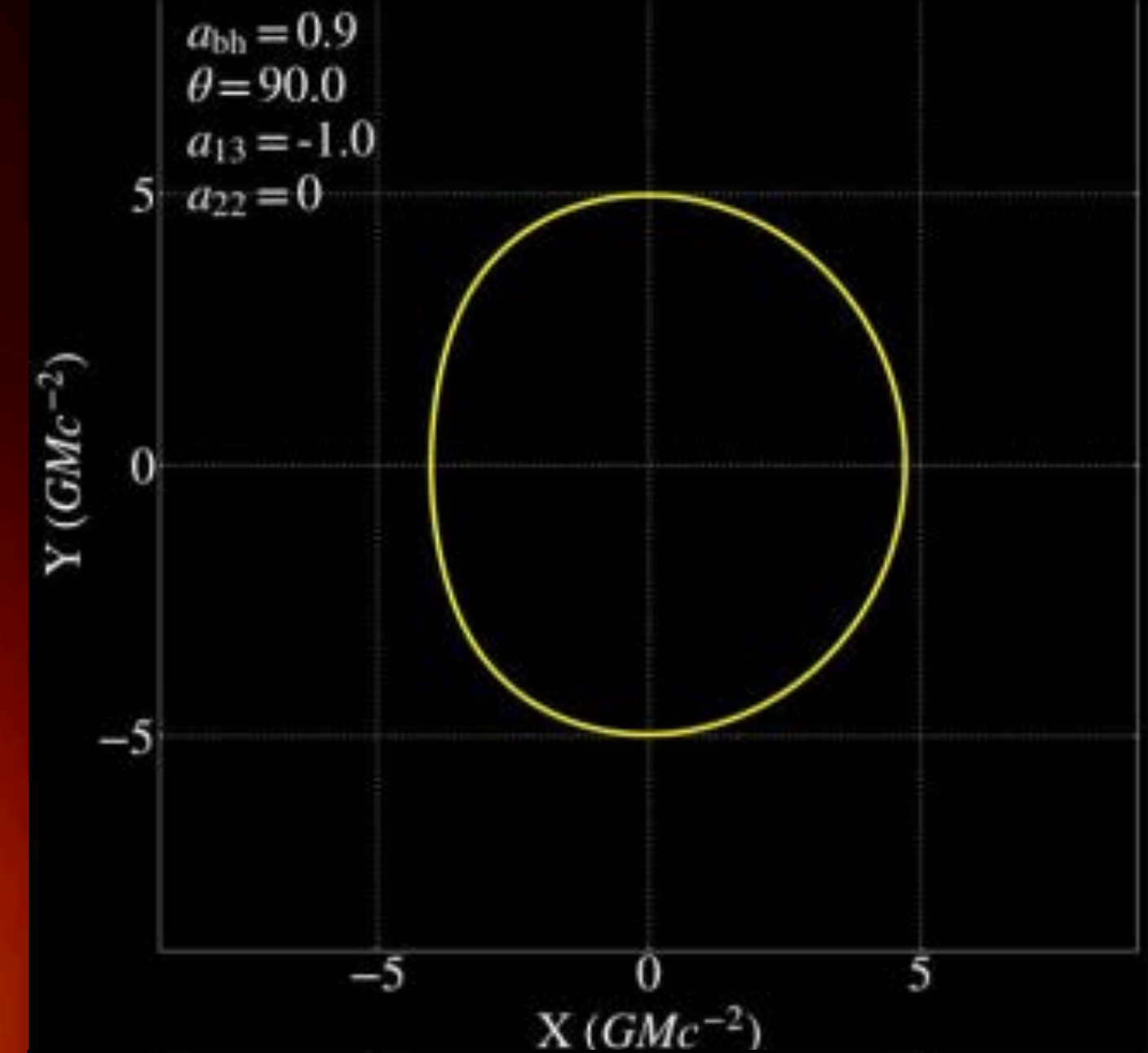
6.5 Billion Solar Mass Black Hole

Deviation from the circle < within 10%
No significant deviations from GR

Black Hole: 4.84-5.2 R_s



Non GR Shadow



Event Horizon Telescope

EHT Collaboration 2019f, ApJL, 875, L6 (Paper VI)

Credit: CK Chan

EHT BLACK HOLE IMAGE
SOURCE: NSF



EHT as a GR Metric Tester

Spherically symmetric space time around a non-spinning black hole

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + r^2d\Omega .$$

Parametrized Post-Newtonian (PPN) formalism

$$-g_{tt} = 1 - \frac{2}{r} + 2 \left(\frac{\bar{\beta} - \bar{\gamma}}{r^2} \right) - 2 \left(\frac{\zeta}{r^3} \right) + \mathcal{O}(r^{-4})$$

General
Relativity
(Schwarzwild metric)

1st order
Deviation
(1PN)

2nd order
Deviation
(2PN)

Weak-field test: $< \sim 10^{-5}$

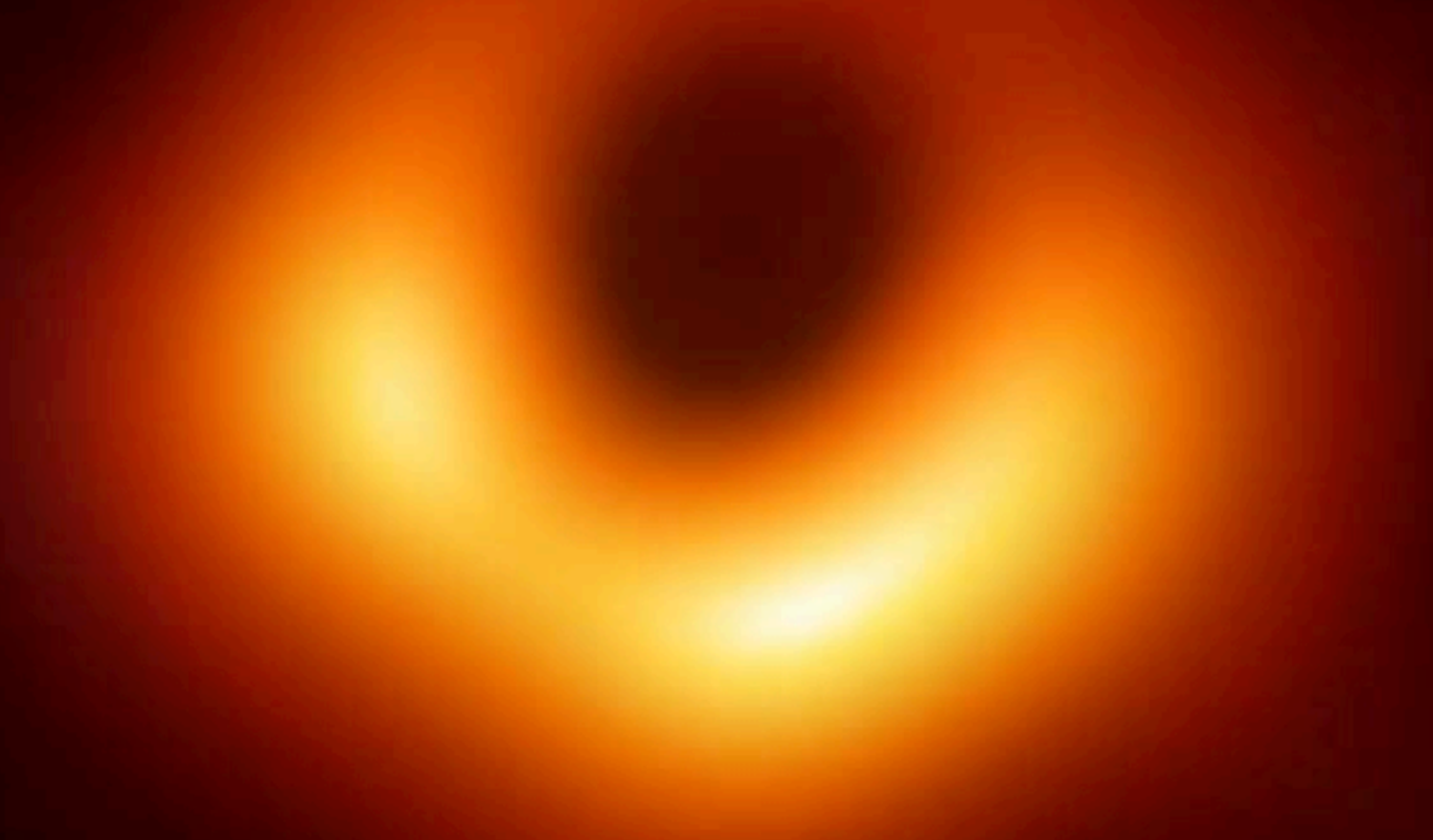
The Diameter of the black hole shadow: sensitive to the 2nd order deviation

$$r_{\text{sh}} = 3\sqrt{3} \left(1 + \frac{1}{9}\zeta \right)$$

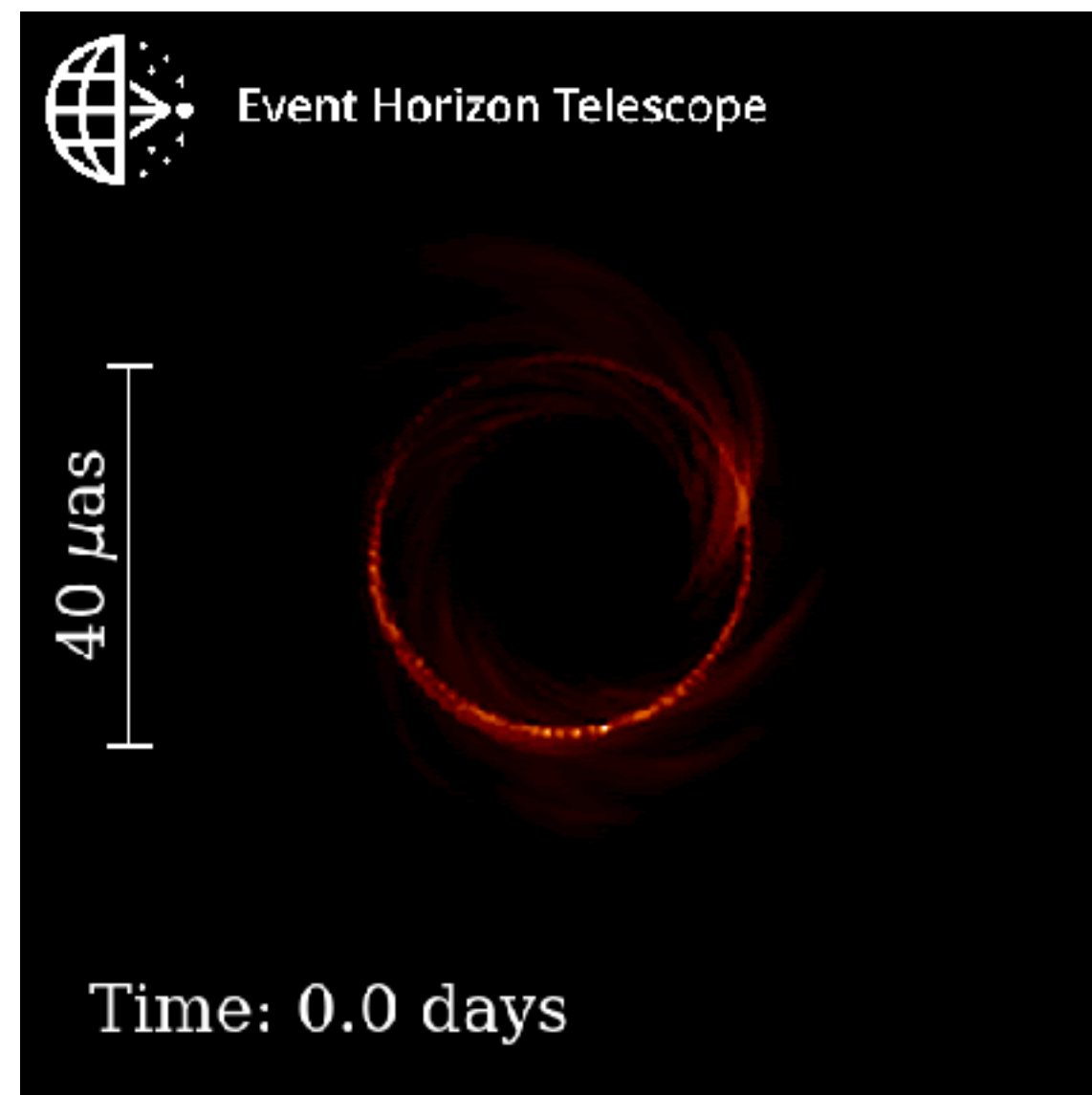
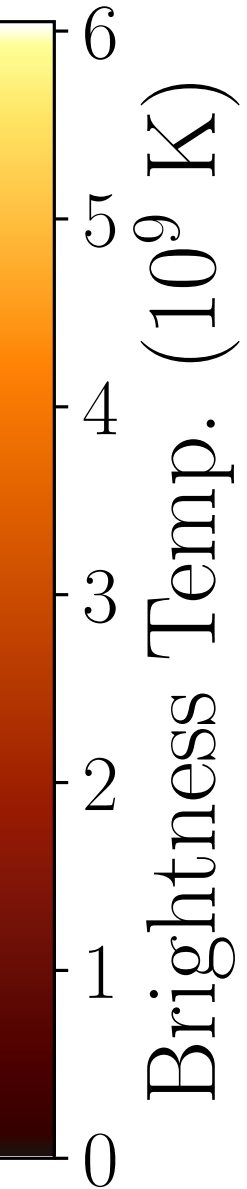
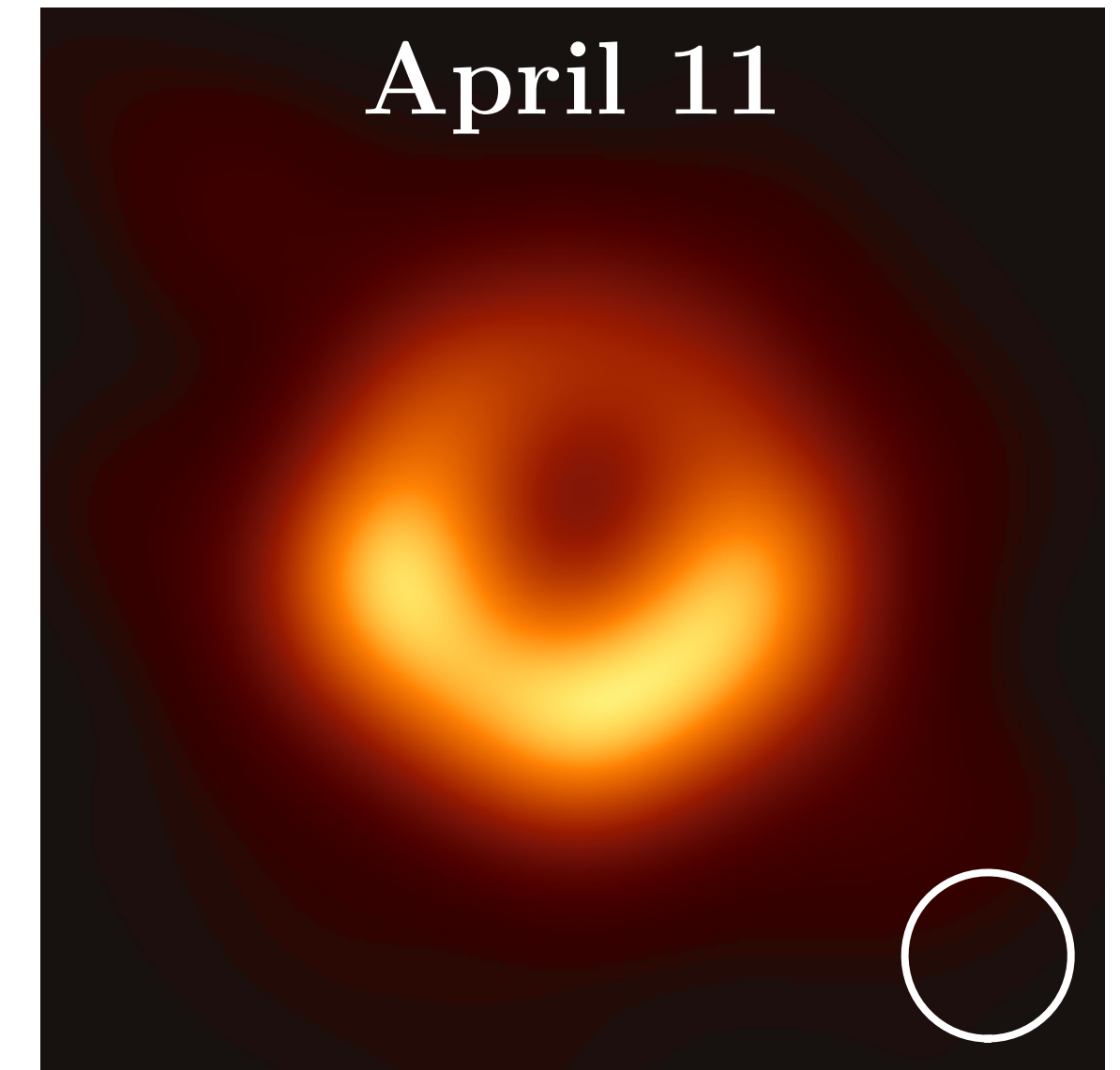
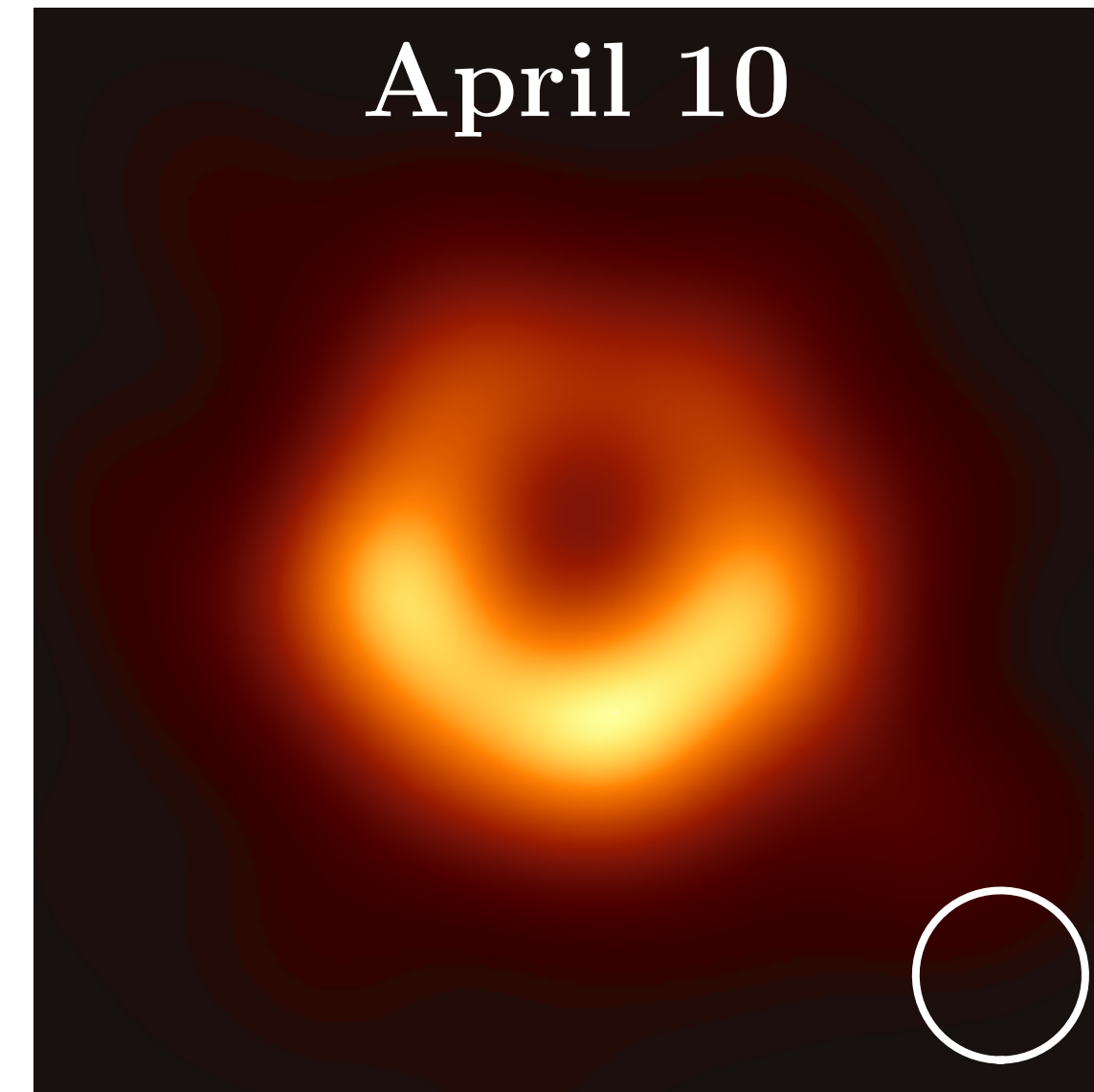
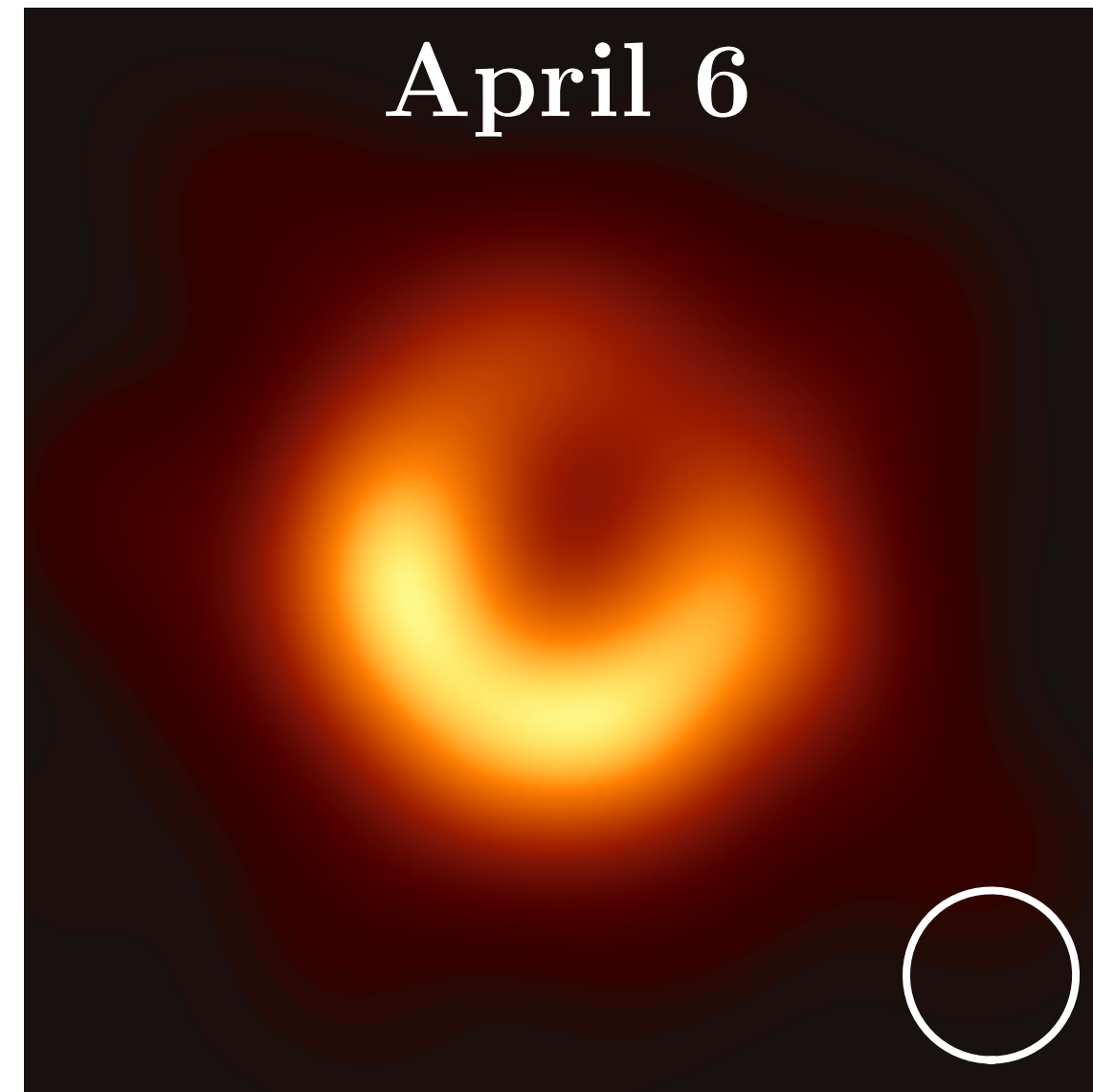
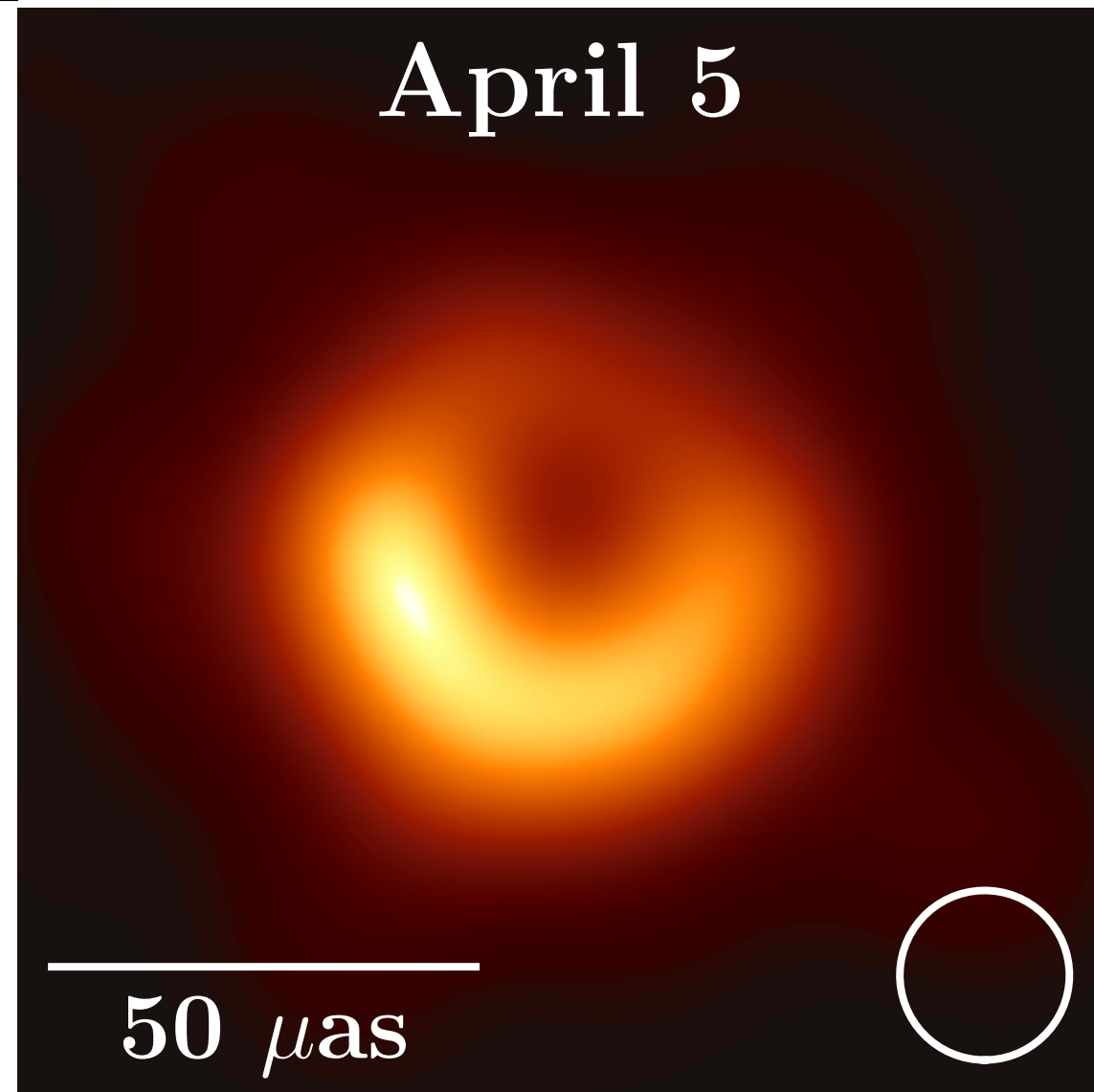
Psaltis et al. 2020, Physical Review Letters



***“EHT measurements of the size of a black hole leads to metric tests
that are inaccessible to weak-field tests”***

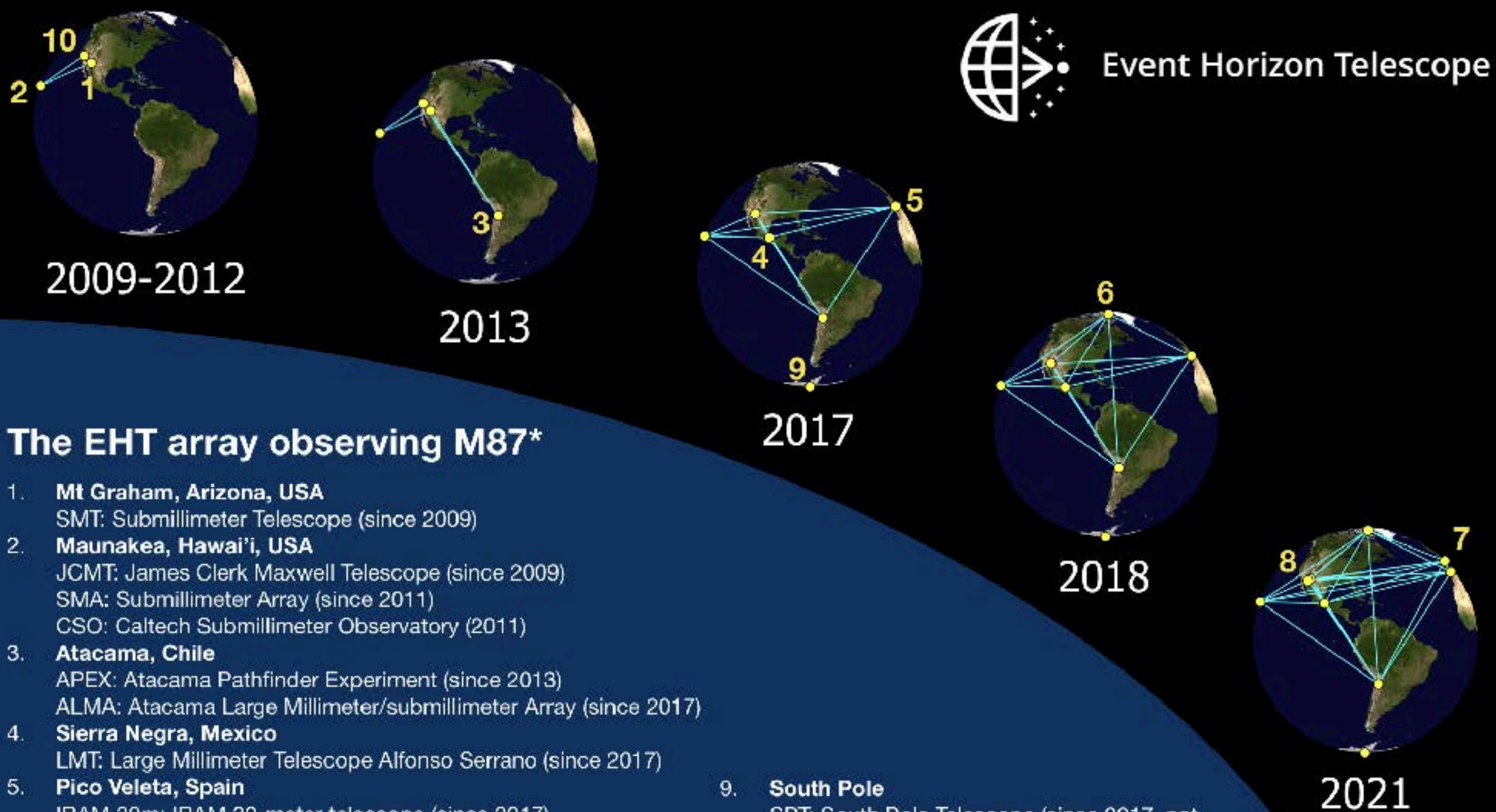


EHT 2017 Images of M87*



Multi-year EHT Observations

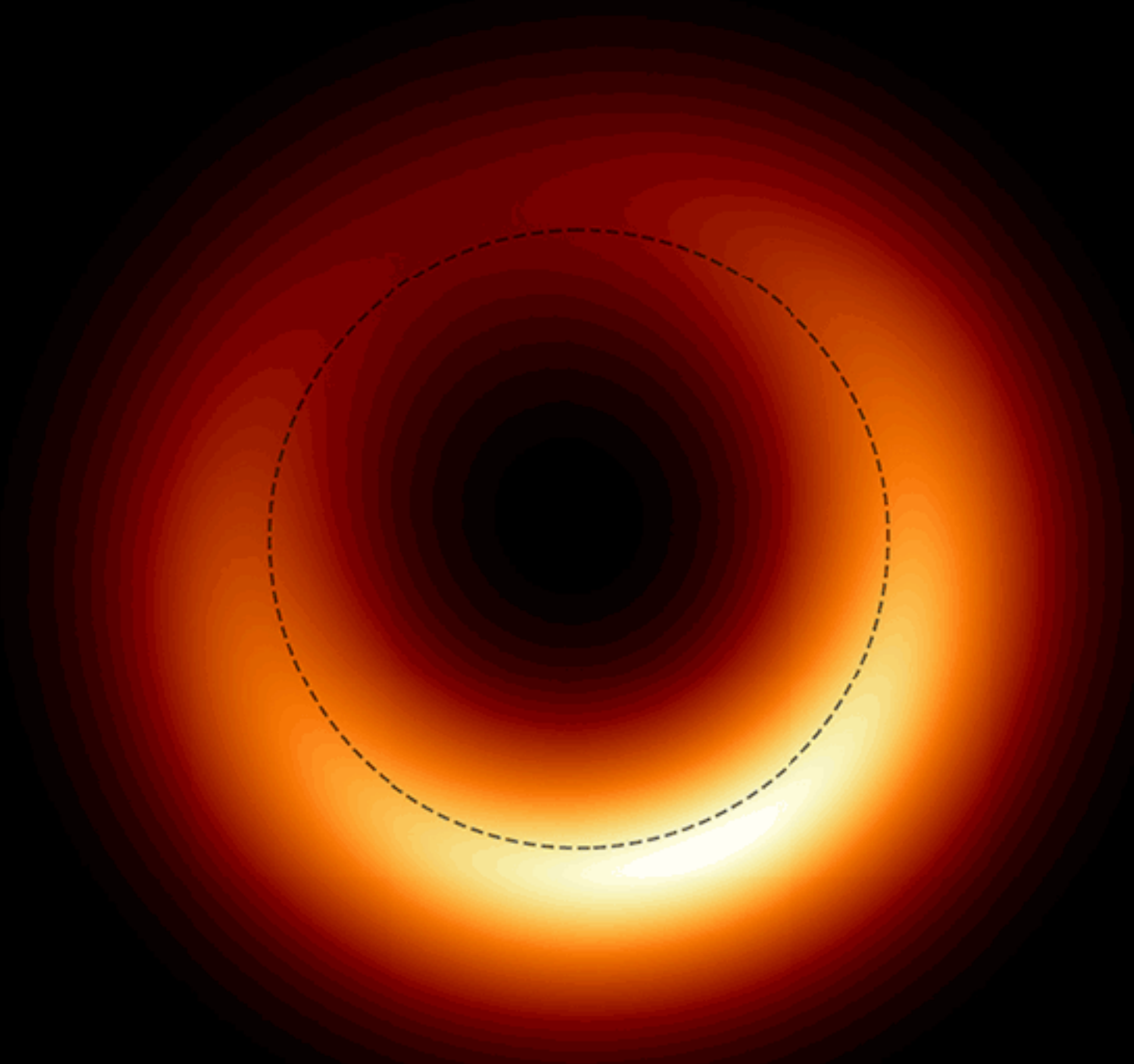
Credit: M. Wielgus, D. Pesce



The EHT array observing M87*

1. **Mt Graham, Arizona, USA**
SMT: Submillimeter Telescope (since 2009)
2. **Maunakea, Hawai'i, USA**
JCMT: James Clerk Maxwell Telescope (since 2009)
SMA: Submillimeter Array (since 2011)
CSO: Caltech Submillimeter Observatory (2011)
3. **Atacama, Chile**
APEX: Atacama Pathfinder Experiment (since 2013)
ALMA: Atacama Large Millimeter/submillimeter Array (since 2017)
4. **Sierra Negra, Mexico**
LMT: Large Millimeter Telescope Alfonso Serrano (since 2017)
5. **Pico Veleta, Spain**
IRAM 30m: IRAM 30-meter telescope (since 2017)
6. **Thule, Greenland**
GLT: Greenland Telescope (since 2018)
7. **Plateau de Bure, France**
NOEMA: NOthern Extended Millimeter Array (since 2021)
8. **Kitt Peak, Arizona, USA**
KP: ARO 12m Radio Telescope (since 2021)
9. **South Pole**
SPT: South Pole Telescope (since 2017, not directly observing M87*)
10. **Cedar Flat, California, USA**
CARMA: Combined Array for Research in Millimeter-wave Astronomy (2009-2013)

2009



©Nature



Wielgus et al. 2020, ApJ

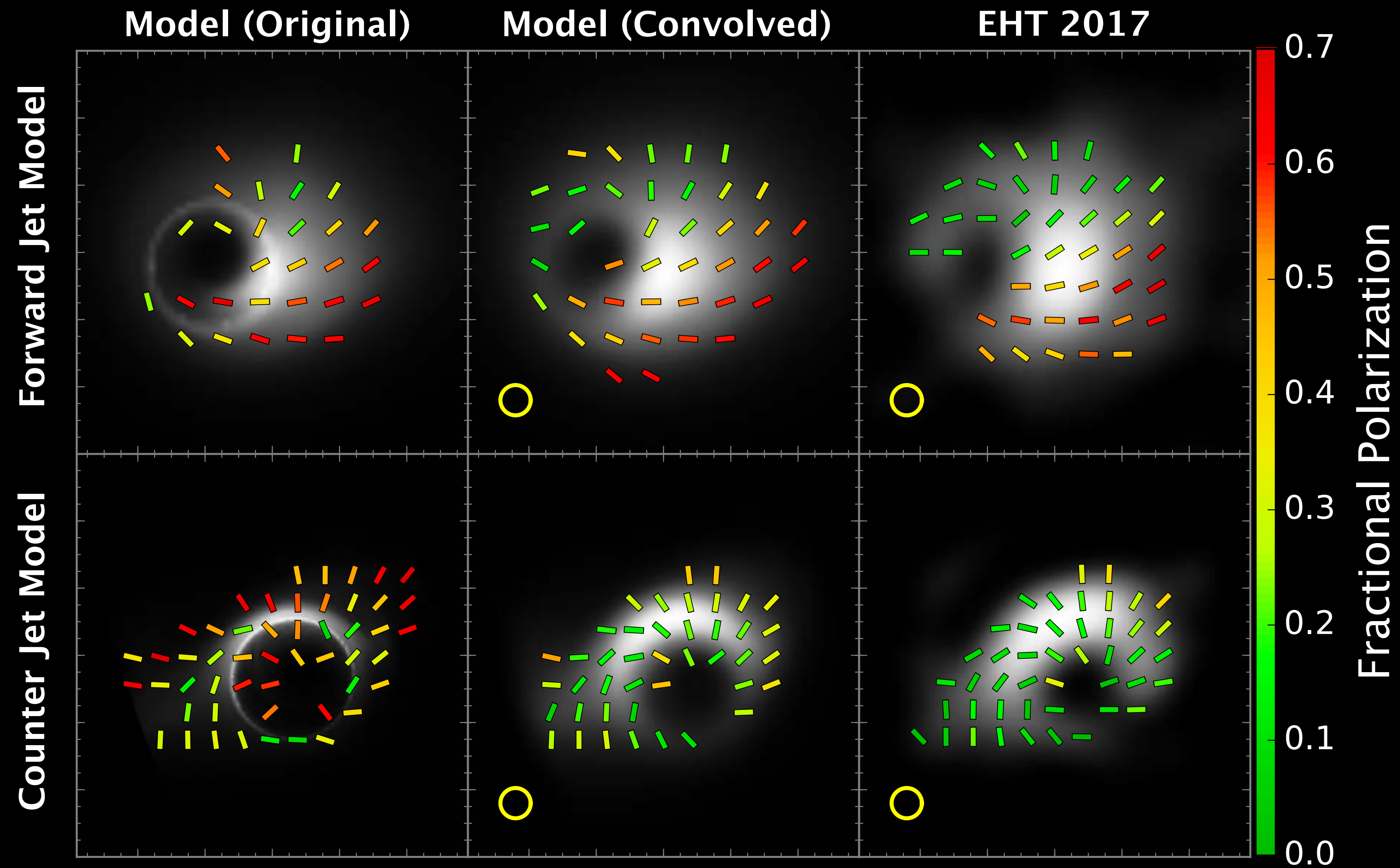
Kazu Akiyama, NERO Symposium 2020, Online, 2020/11/17 (Tue)

First M87 Results: Where are we now?

- Einstein's GR has passed a new test at an extremely strong gravitational field
- The strongest evidence for the presence of a supermassive black hole
- The M87 central black hole is most likely spinning
- An AGN and associated jet are powered by a supermassive black hole
- The stellar dynamical mass is correct (6.5 billion masses)
- Day-to-day & Multi-year variations on horizon scale

Dawn of a New Era of Black Hole Astrophysics

M87 Polarimetry: Coming Soon



Akiyama et al. 2017, Chael et al. 2016

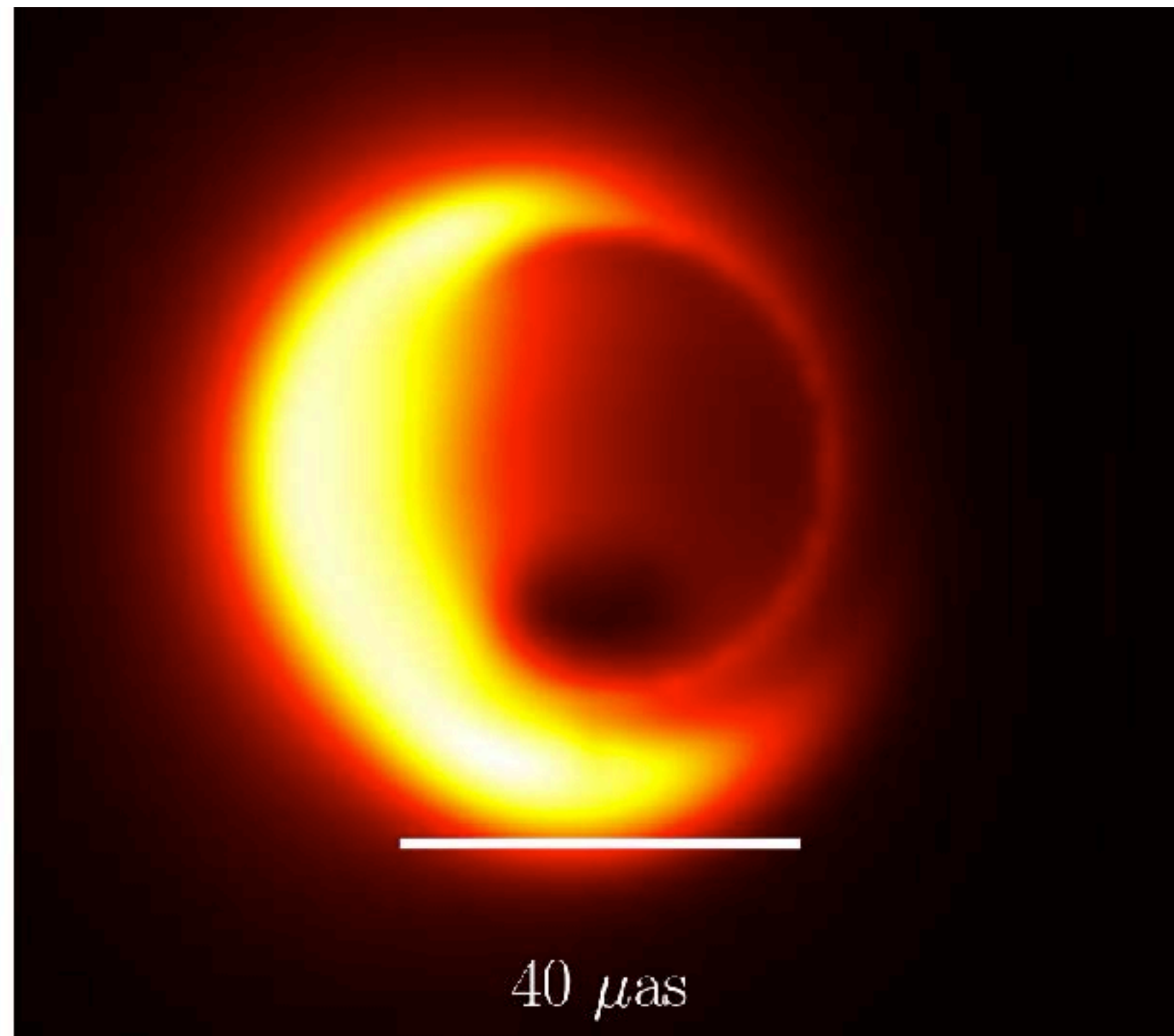
Kazu Akiyama, NEROC Symposium 2020, Online, 2020/11/17 (Tue)



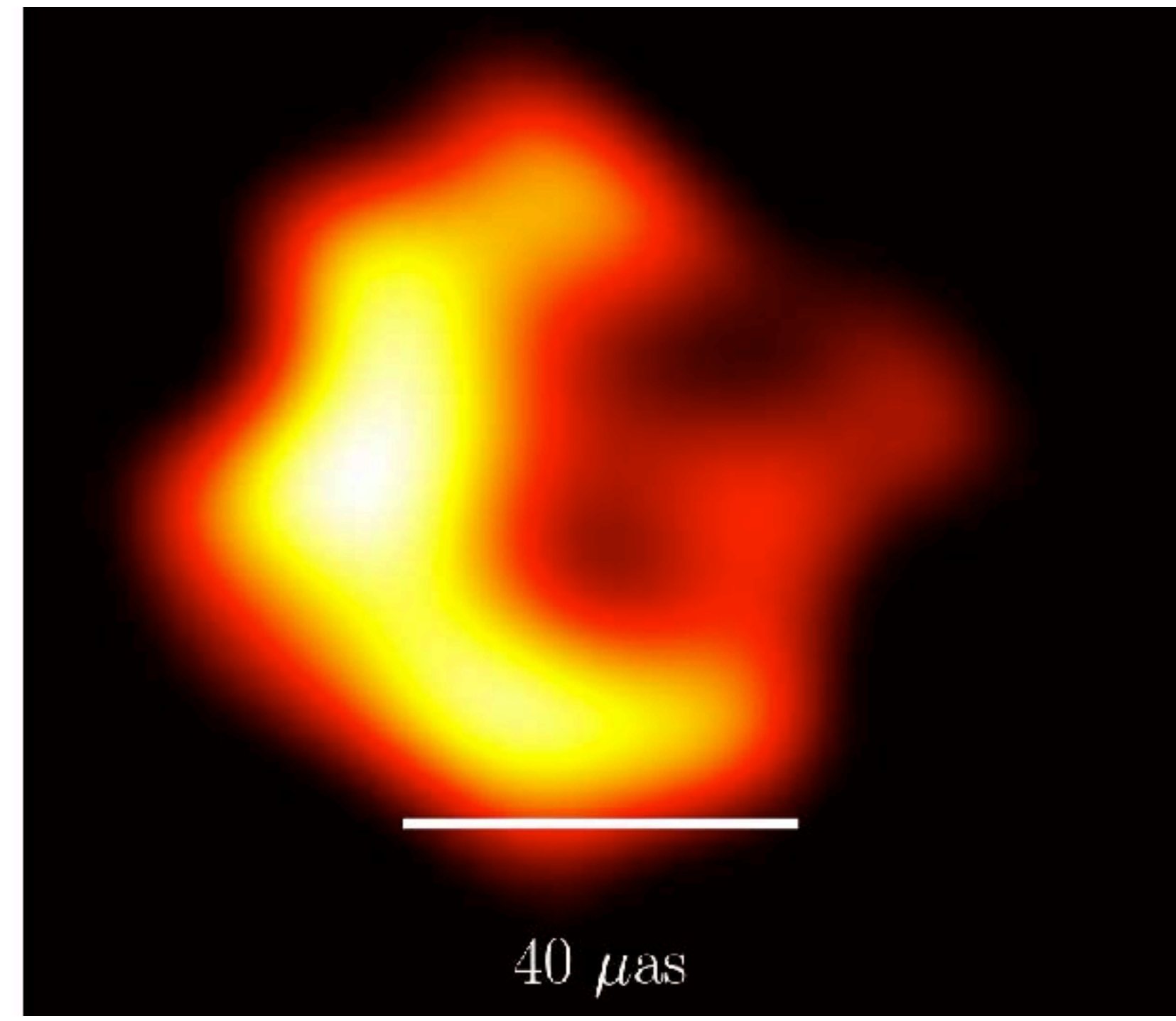
First Horizon-scale Imaging of Sgr A*: Coming soon

RML Video Reconstruction (3D Imaging)

Model

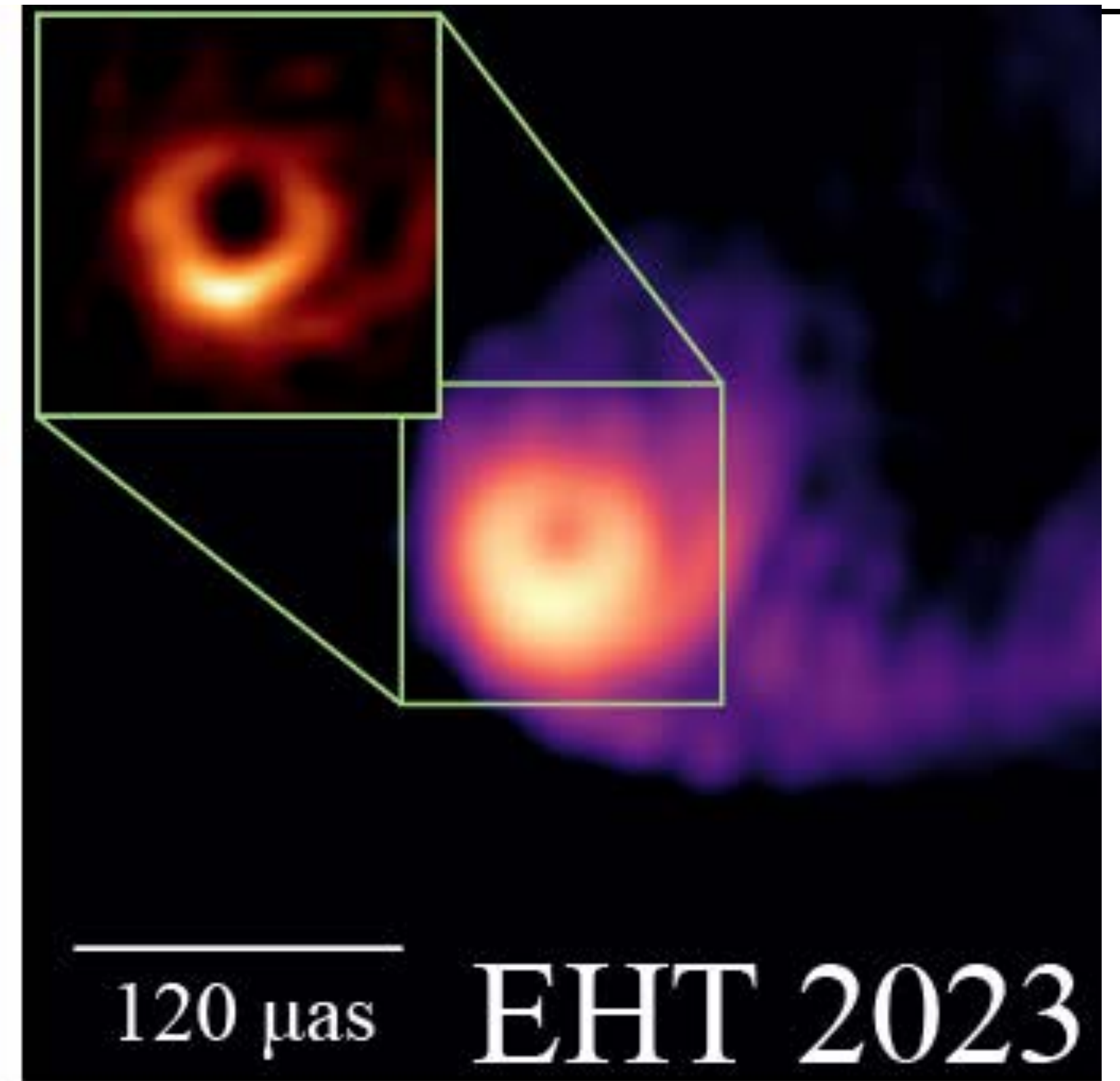
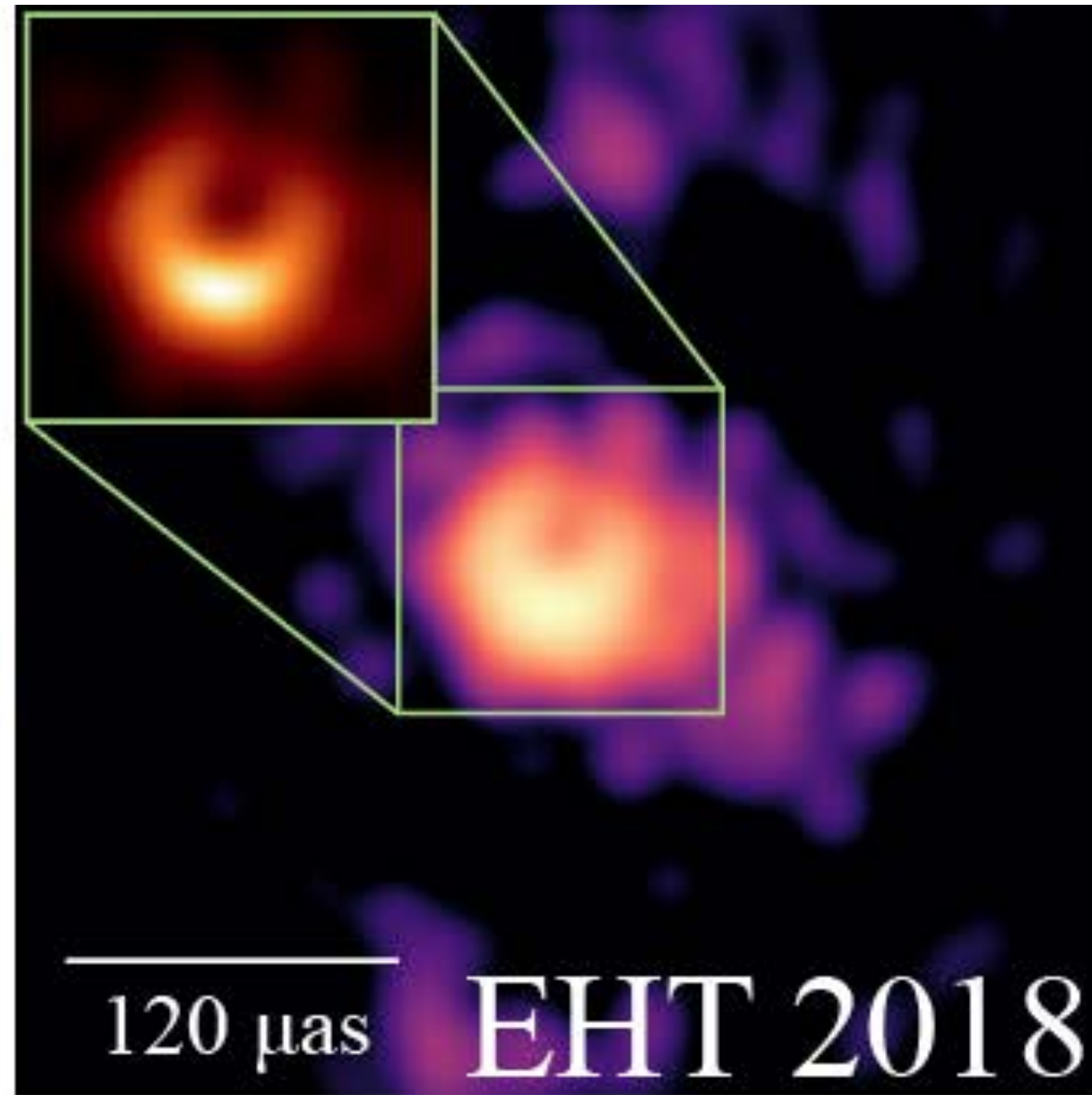
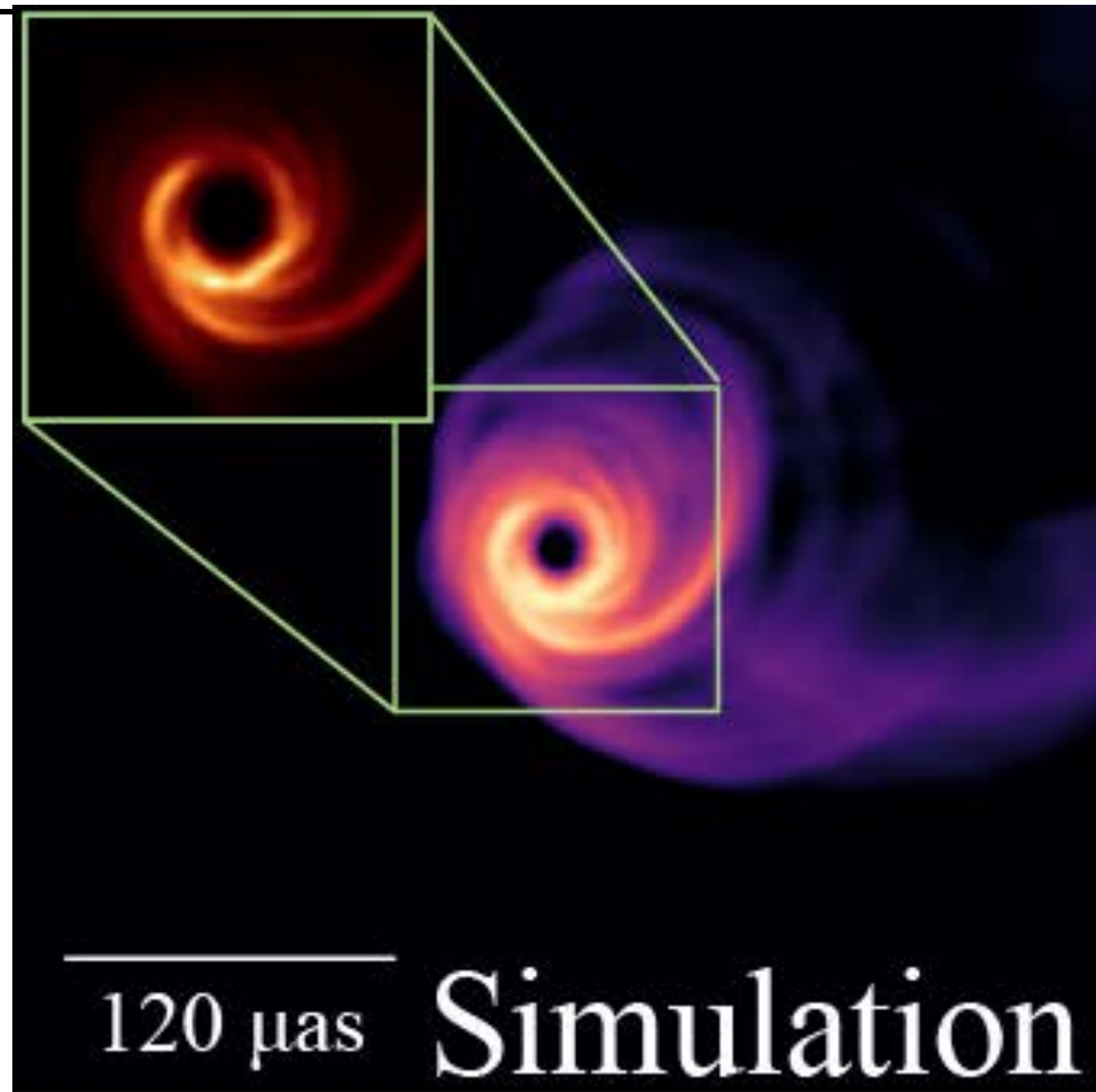


EHT 2017 Simulation



Model: Broderick & Loeb 2006, Imaging simulation: Kotaro Moriyama

EHT 2020s: Deeper, Sharper & Multi-frequency Images



230 GHz
EHT2017+GLT

230+345 GHz
EHT2017+GLT+KP+NOEMA+OVRO

Further new capabilities: Faraday Rotation Imaging (5 Dimensional Imaging)

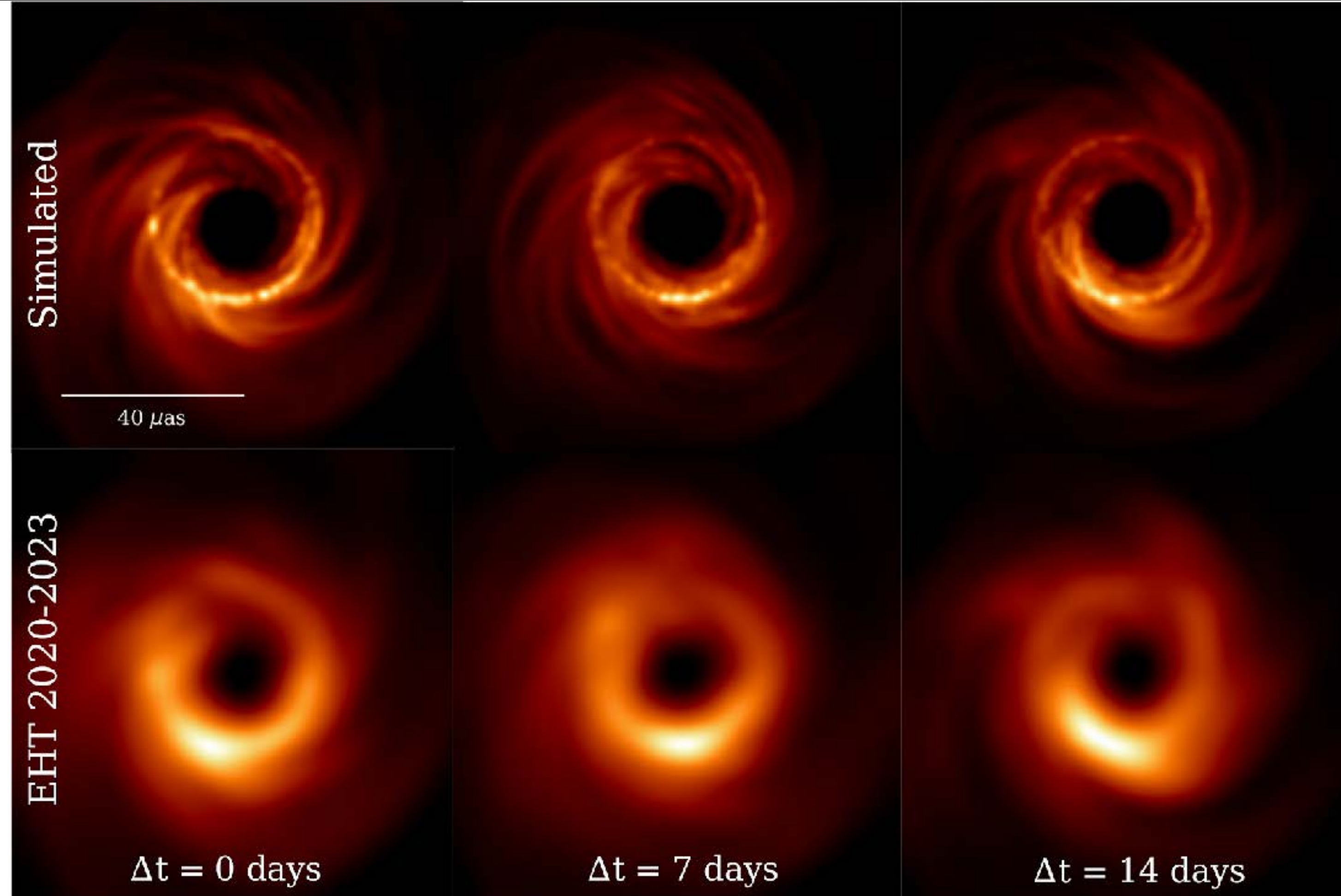
Active developments of Multi-scale, multi-dimensional imaging techniques



Simulations: Andrew Chael, Imaging: Kazu Akiyama

Kazu Akiyama, NEROC Symposium 2020, Online, 2020/11/17 (Tue)

EHT 2020s: Tracking Evolving Features

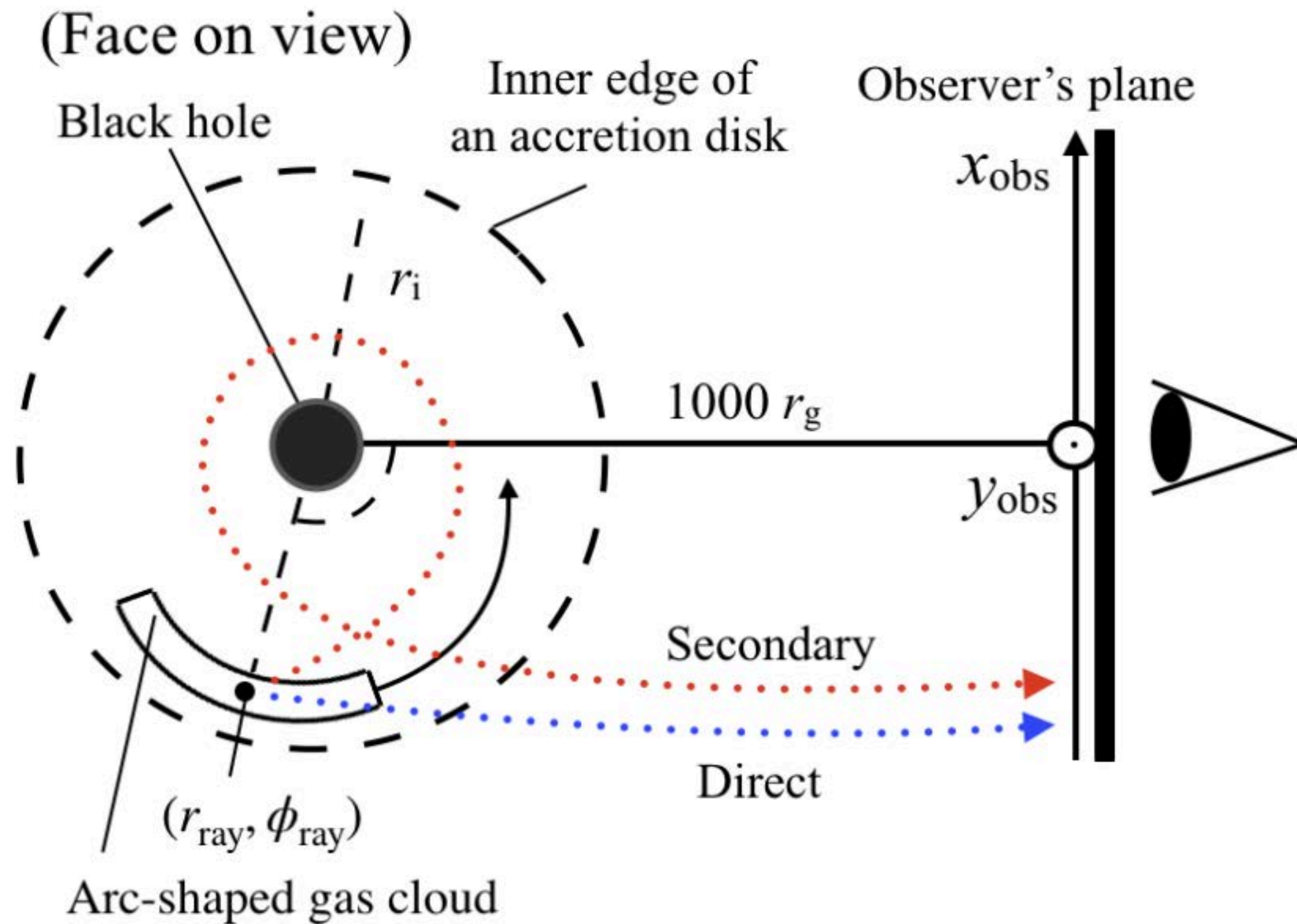


Regular monitoring observation capabilities on weekly scales

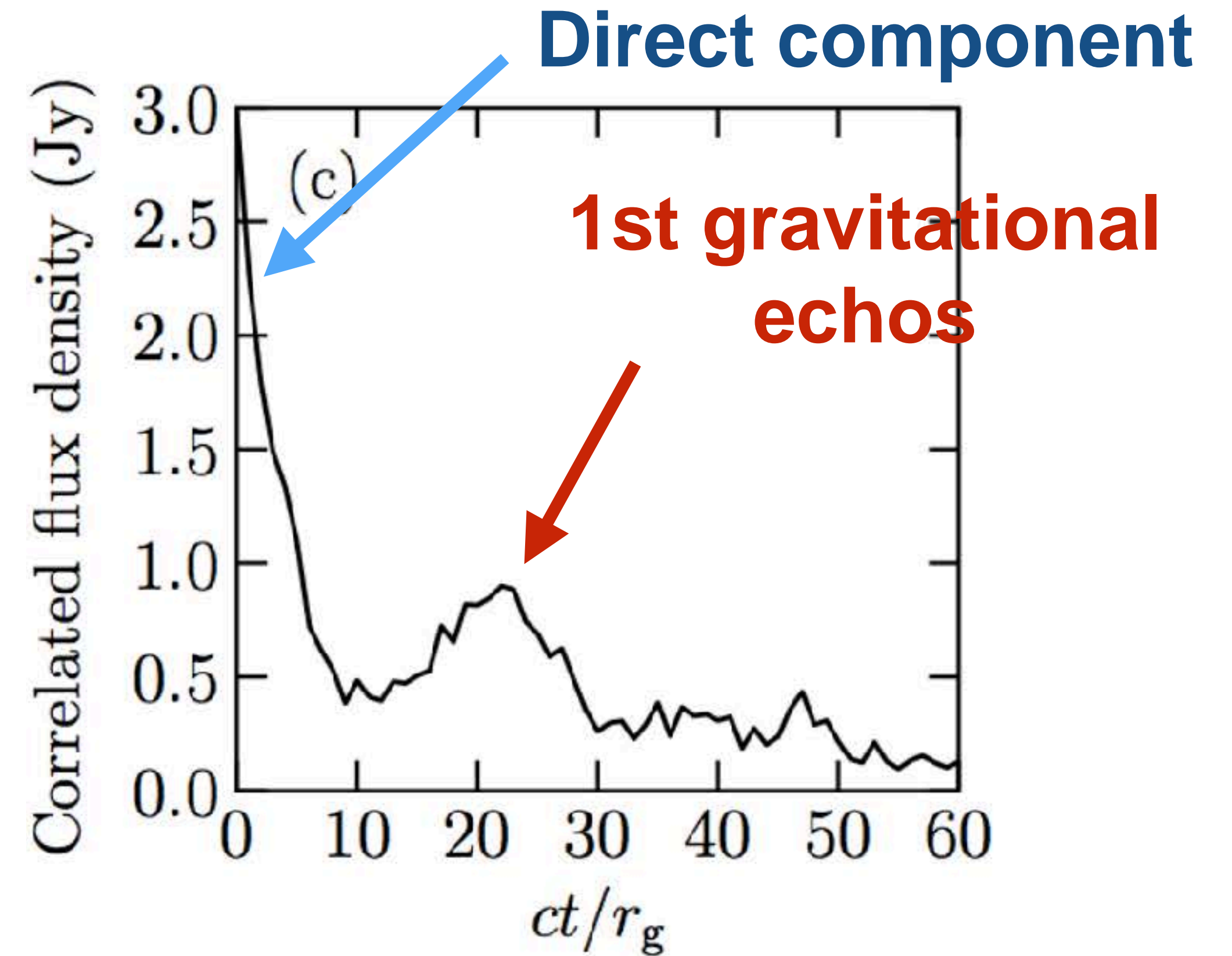
Simulations: Charles Gammie, George Wong et al., Imaging: Michael Johnson

Kazu Akiyama, NERO Symposium 2020, Online, 2020/11/17 (Tue)

EHT2020: Spatially Resolved Time-domain Signature of Echos

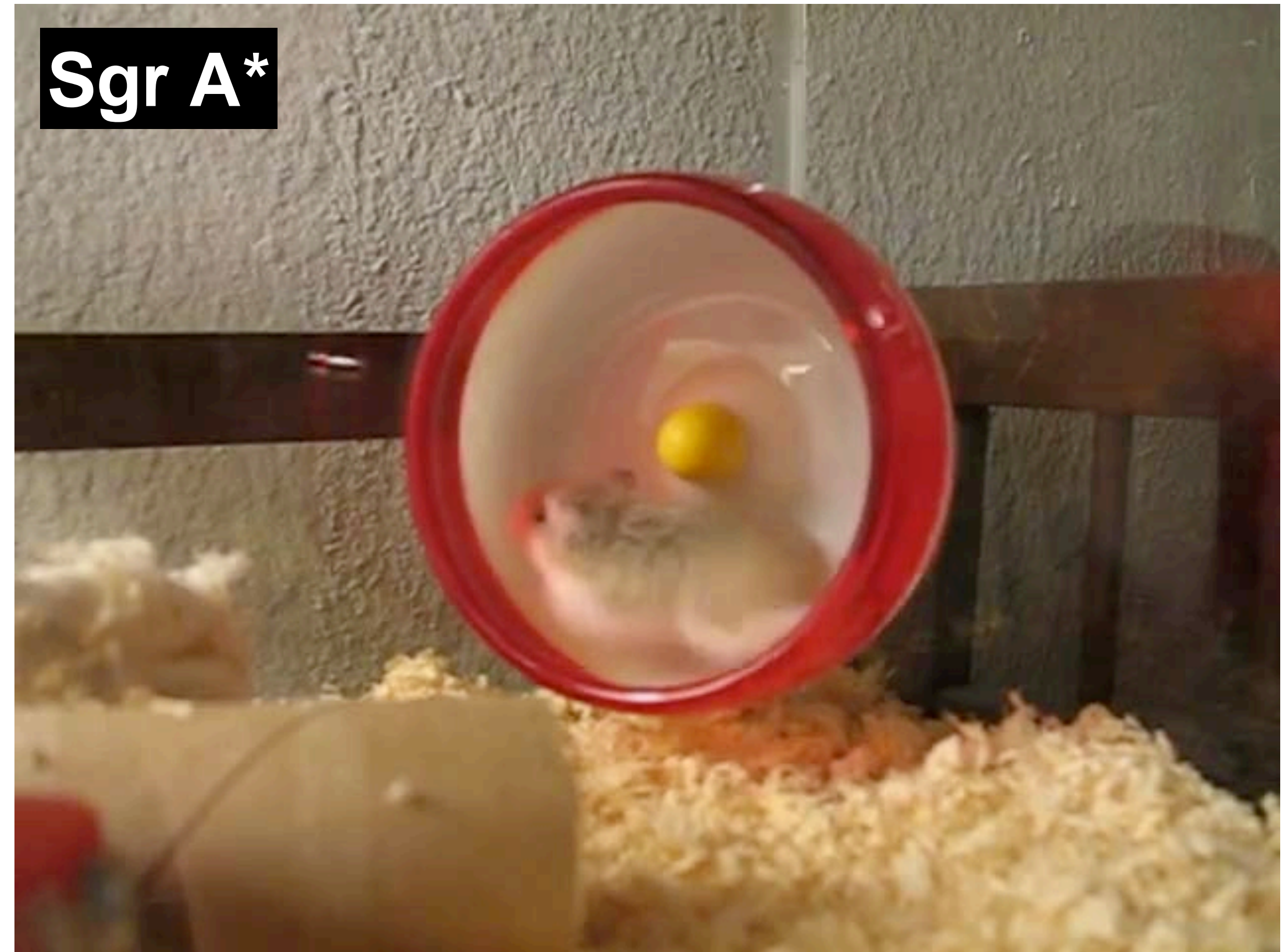
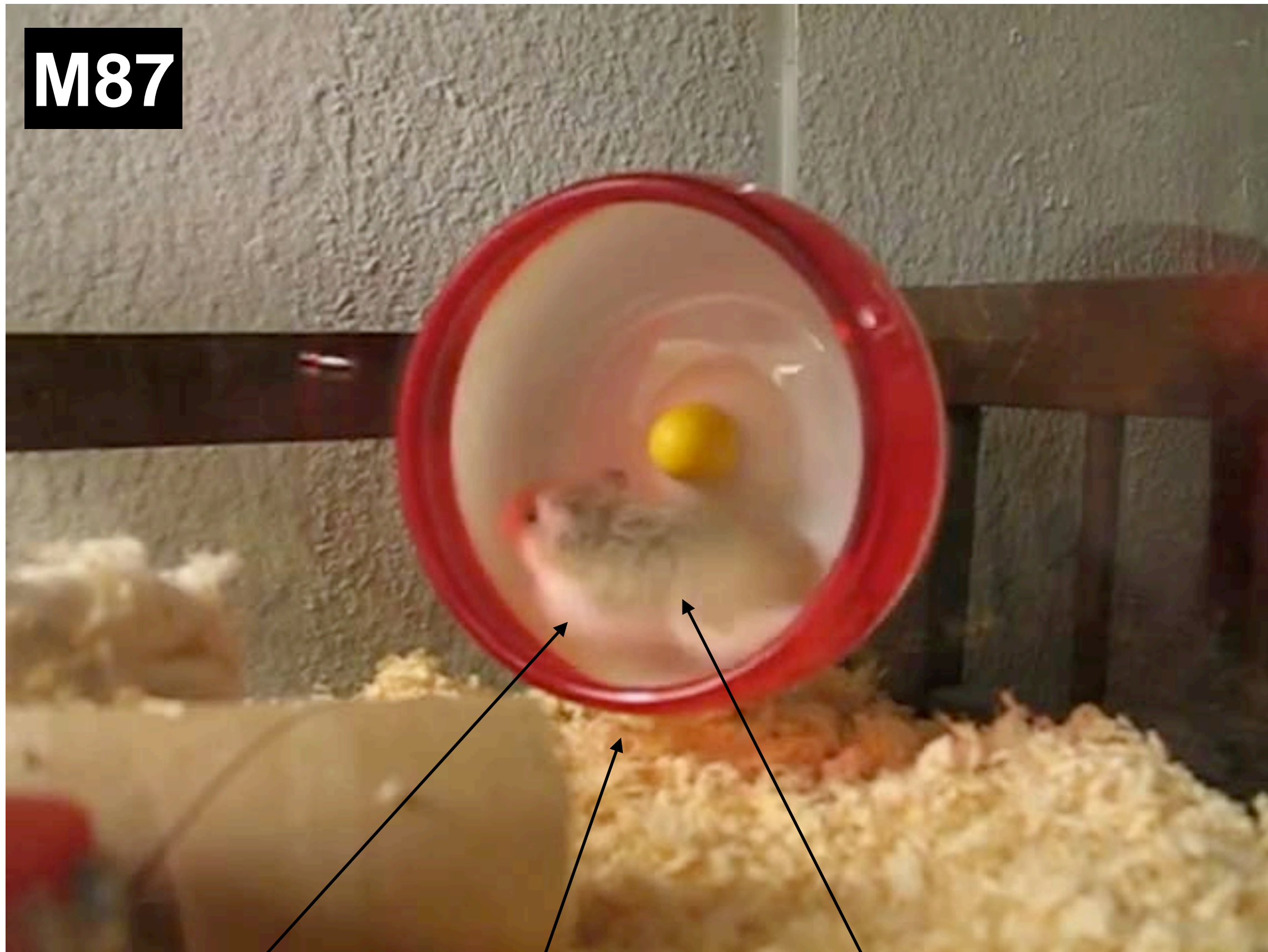


Stacked profile of correlated flux



See Kotaro Moriyama (MIT)'s Poster

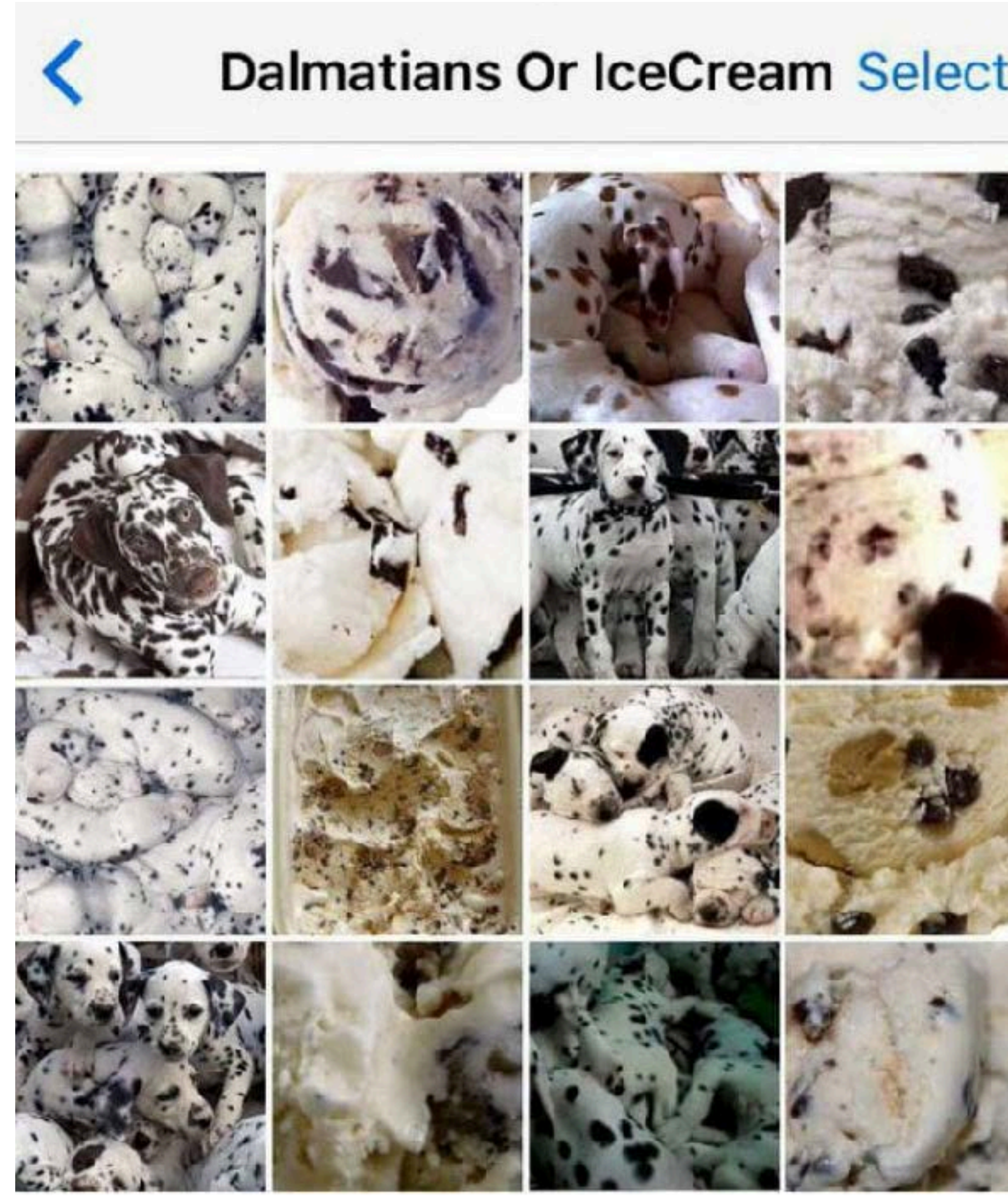
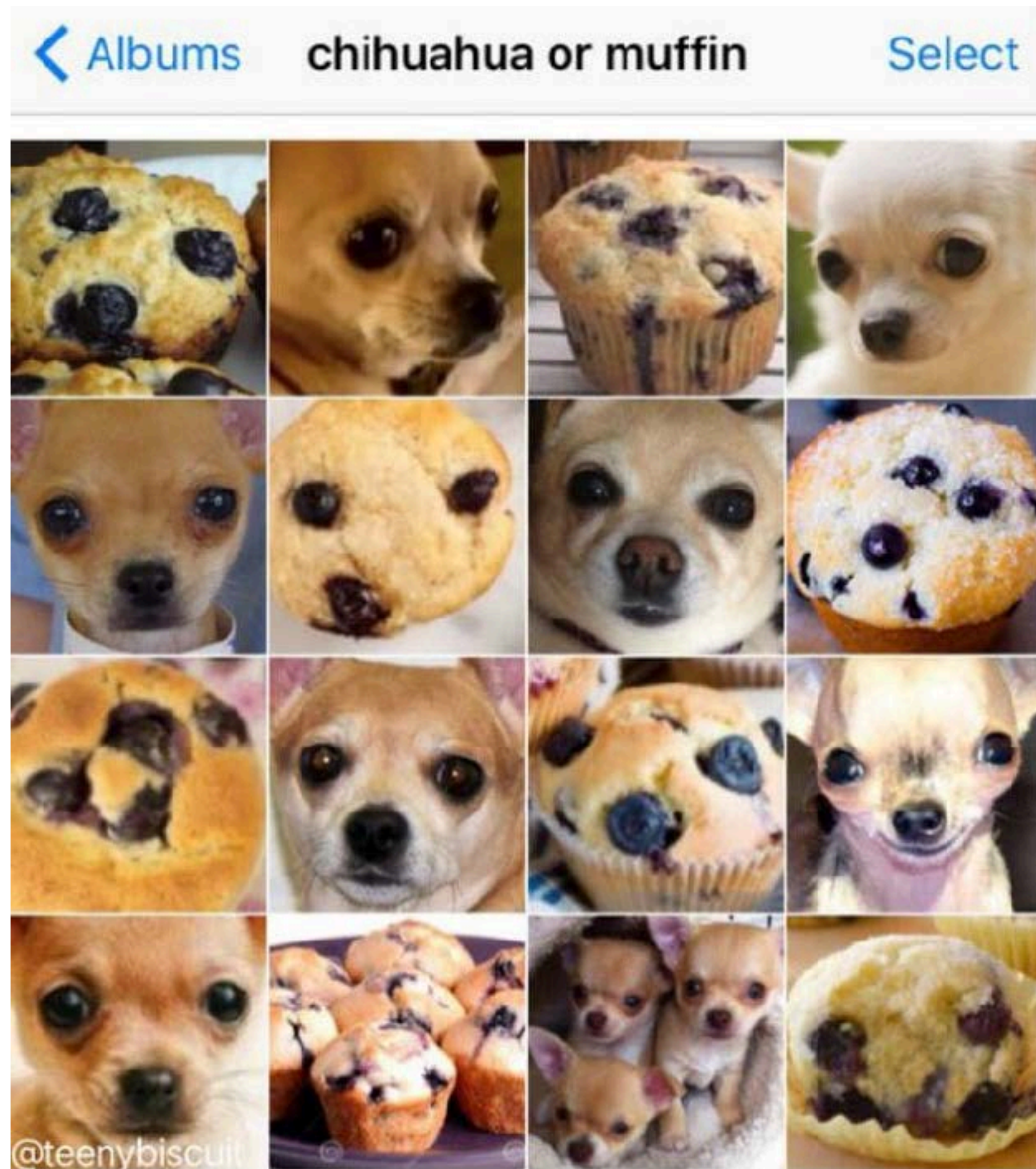
EHT 2020s: Precision Black Hole Astrophysics



Mass ? Spin ? Accretion flow types?
Viewing Geometry?

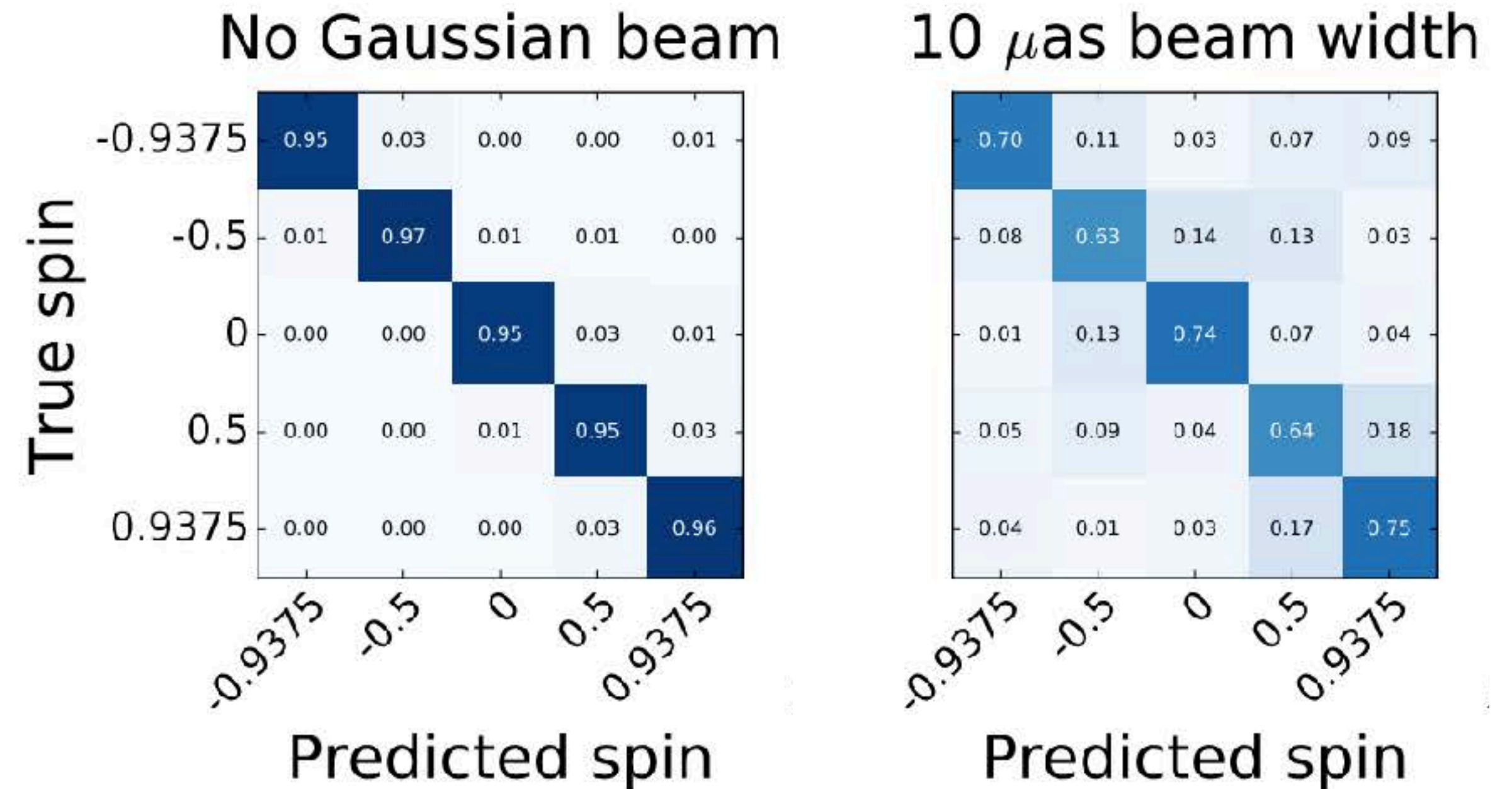
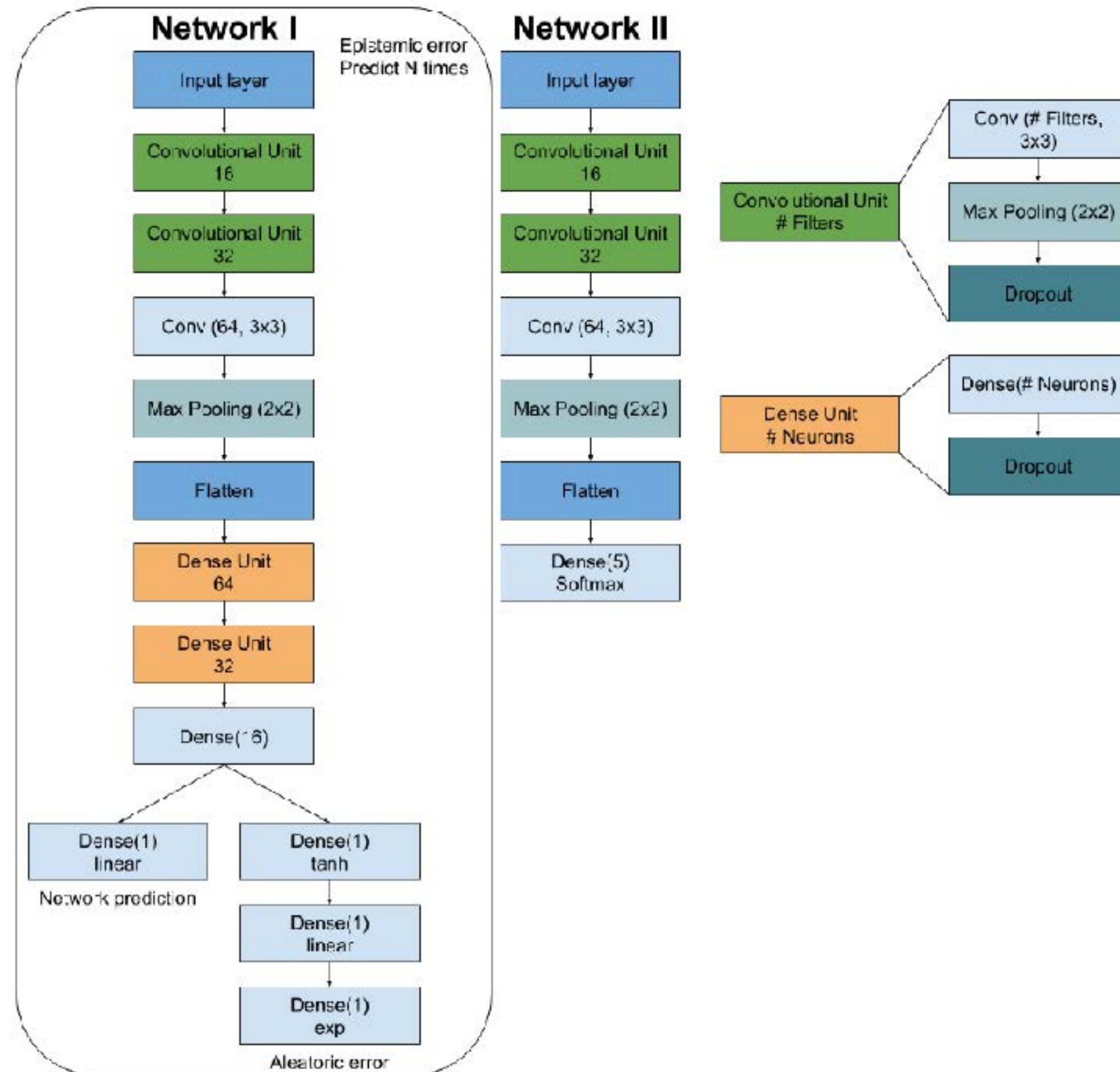
EHT 2020s: Precision Black Hole Astrophysics

Are black hole images confusing for scientists and/or AI?



EHT 2020s: Precision Black Hole Astrophysics

Current forecast: Horizon-scale images are much less confusing!

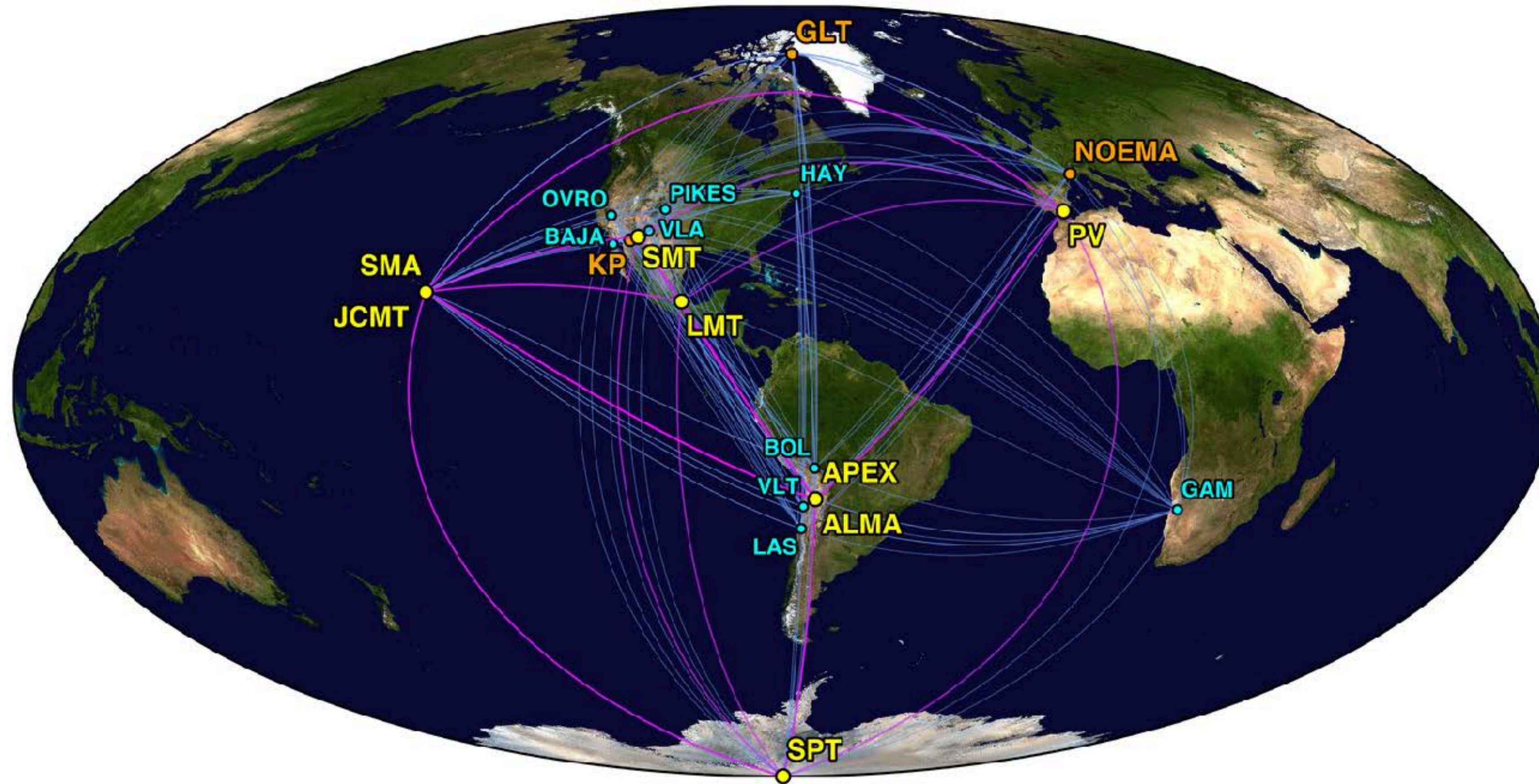


Van der Gucht et al. 2019

Van der Gucht+, Yao-Yu Lin+, Sun & Bouman+



next generation Event Horizon Telescope

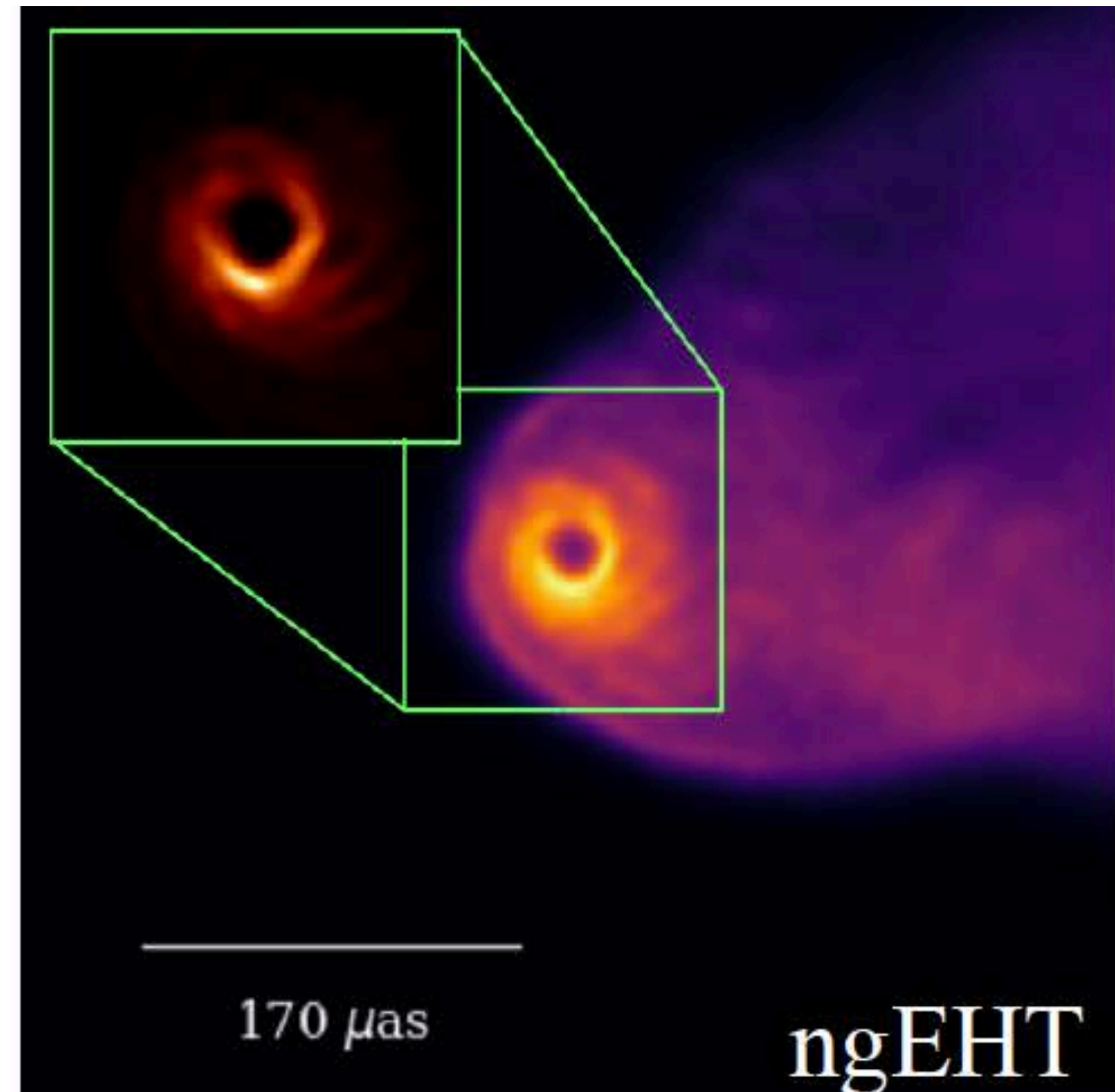
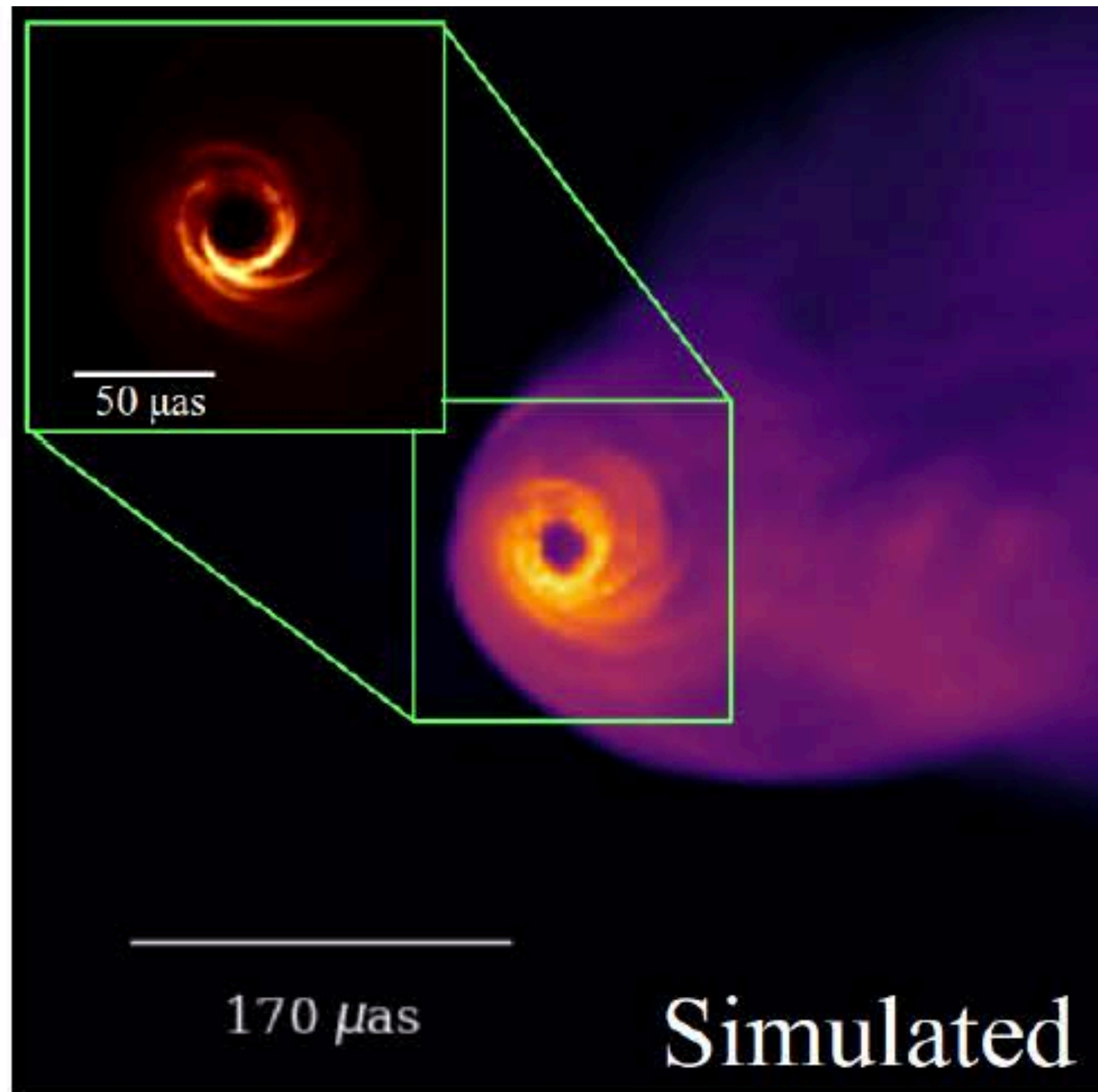


Phase I: 2019-2023 (Array Design Phase + MIT Haystack as a potential new site)

Phase II: 2023- (Constructions of several new sites)



M87 ngEHT images: ~10 years from now



Blackburn et al. 2019; Doeleman et al. 2019 (Astro2020 Decadal Survey White Papers)



Thanks for listening!

